

# Orbital

Space where electrons are found

Orbitals are characterized by set of quantum numbers (n,l,m)

Solution of Schrodinger wave funct<sup>n</sup> {

<p>Principal Quantum no.</p> <p>Azimuthal <math>\rightarrow</math> —</p> <p>Magnetic <math>\rightarrow</math> —</p>	<p><math>n \rightarrow</math> shell no.</p> <p><math>l \rightarrow</math> Subshell</p> <p><math>m \rightarrow</math> orbital</p>
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# Quantum Number

Set of numbers which represent the position of electron in an atom

Concept came from **Quantum Mechanics**

Schrodinger Eqn

$$\hat{H}(\psi) = E(\psi)$$

Solution of this Equation gave three set of quantum numbers

$$\frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} + \frac{\partial^2 \psi}{\partial z^2} + \frac{8\pi^2 m}{h^2} (E - V) \psi = 0$$

$\begin{cases} n \\ l \\ m \end{cases}$

## Schrodinger Eqn

$$\hat{H}(\psi) = E(\psi)$$

Solution of this Equation gave **three set of quantum numbers**

Schrodinger {

**Principal quantum numbers(n)**

**Magnetic quantum numbers(m)**

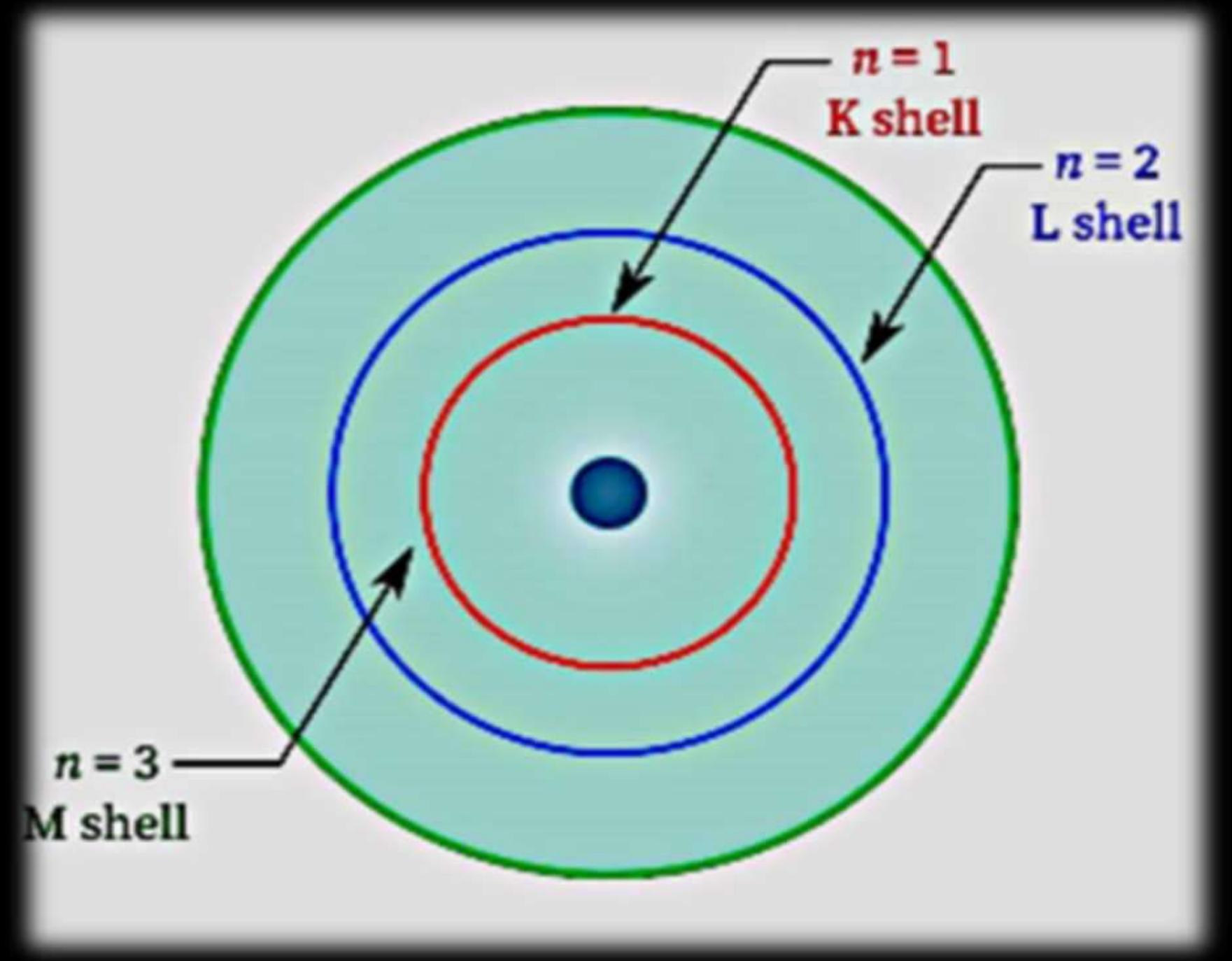
**Azimuthal quantum numbers(l)**

\* But 4<sup>th</sup> Quantum no.

Came from  
experiment

**Spin quantum numbers(s)**

# Principal quantum numbers(n)



shell no.	(2)	(3)
Subshells	(2)	(3)
orbitals	(4)	(9)
$e^-$	(8)	(18)

\* size ( $n < 3$ )

\* energy ( $n < 3$ )

- \* It represents the shell no.
- \* Also tell about size & energy of shell as  $(n) \uparrow \Rightarrow \text{size} \uparrow \text{ energy} \uparrow$

- \*  $(n)$  shell no.

$$\left\{ \begin{array}{l} \rightarrow \text{subshell} = (n) \\ \rightarrow \text{orbitals} = (n^2) \\ \rightarrow e^- = (2n^2) \end{array} \right.$$

# Azimuthal (Angular momentum) quantum numbers(I)

(\*) Shape of orbital in given subshell

Ex:  $n = 1 \Rightarrow l = 0 \Rightarrow 1s$

(\*) Name of subshell

$n = 2 \Rightarrow l = 0, 1 \Rightarrow 2s, 2p$

(\*) Its value (0 to  $(n-1)$ )

(\*)  $l = 0 \Rightarrow s$ -subshell  $\Rightarrow$  spherical shape



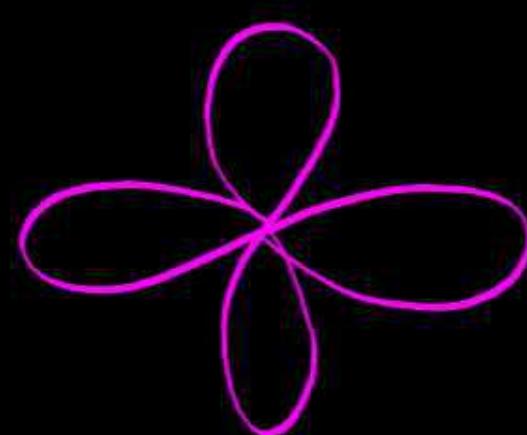
$n = 3 \Rightarrow l = 0, 1, 2$

3s, 3p, 3d

$l = 1 \Rightarrow p$ -subshell  $\Rightarrow$  dumbbell shape



$l = 2 \Rightarrow d$ -subshell  $\Rightarrow$  double dumbbell shape



$l = 3 \Rightarrow f$ -subshell  $\Rightarrow$  not defined

(\*) Orbital angular momentum =  $\sqrt{l(l+1)} \frac{h}{2\pi}$

(\*) No. of orbitals =  $(2l+1)$       (\*) no. of e<sup>-</sup> =  $2(2l+1) \Rightarrow 4l+2$

NOTE:

1) Bigger ( $n$ )  $\leftrightarrow$  far more energy

2) Same ( $n$ )

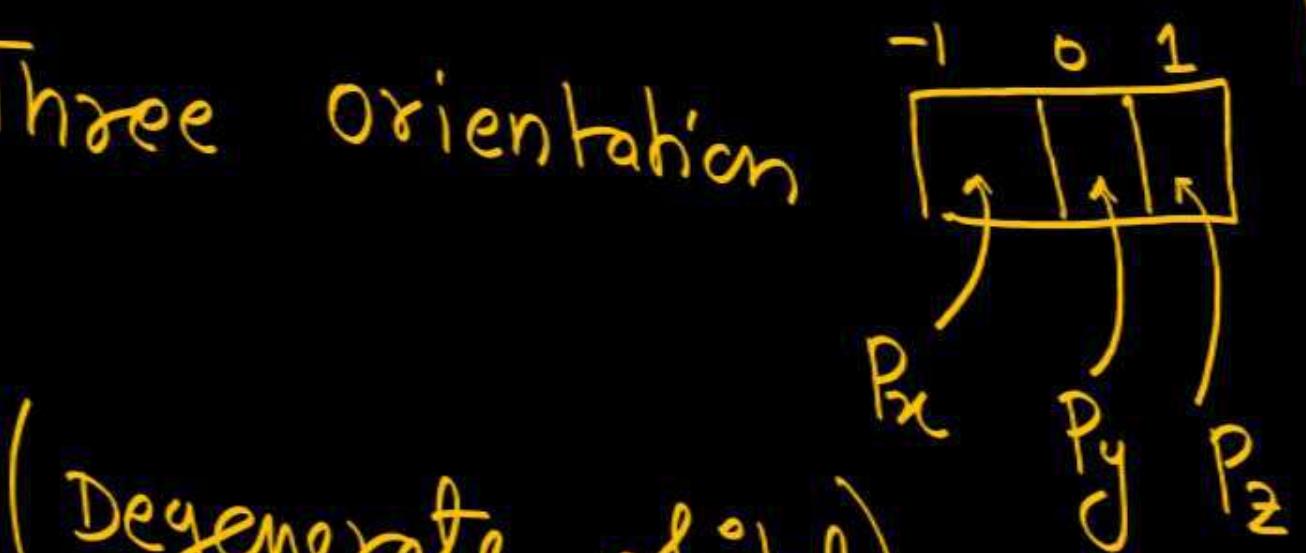
s > p > d > f

Penetration power

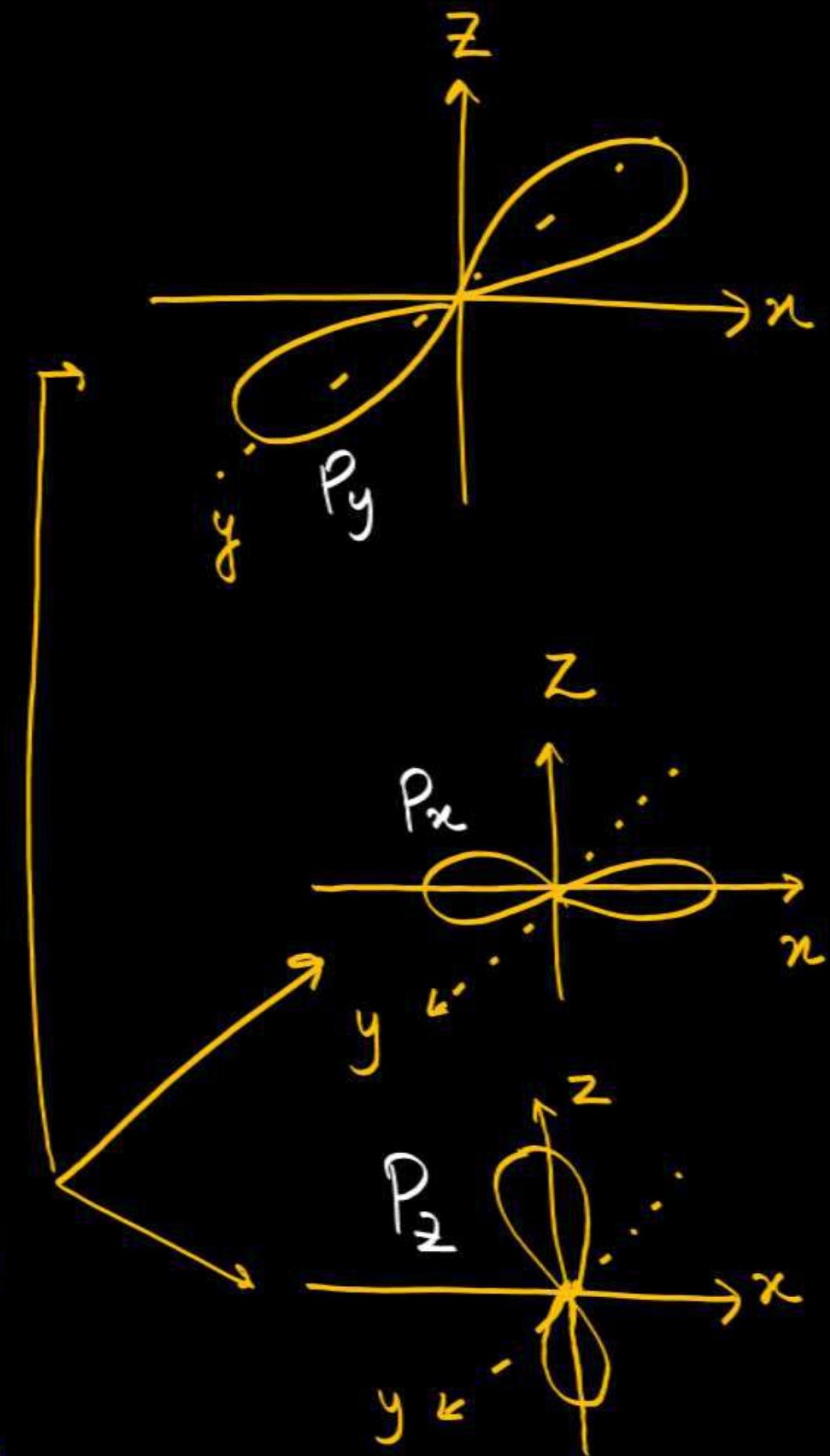
3) Energy  $\Rightarrow$  s < p < d < f

## Magnetic quantum numbers ( $m_l$ )

- \* (3-d) spatial arrangement of orbitals
- \* Its value from  $(-l)$  to  $(+l)$  → Total values  $(2l+1)$
- \* Each orbital have diff orientation
- \* Total orbitals =  $(2l+1)$   
*(orientation)*
- \*  $l=0 \Rightarrow (s) \Rightarrow m=0 \Rightarrow$  one orientation  $\square$
- \*  $l=1 \Rightarrow (p) \Rightarrow m = -1, 0, 1 \Rightarrow$  Three orientation

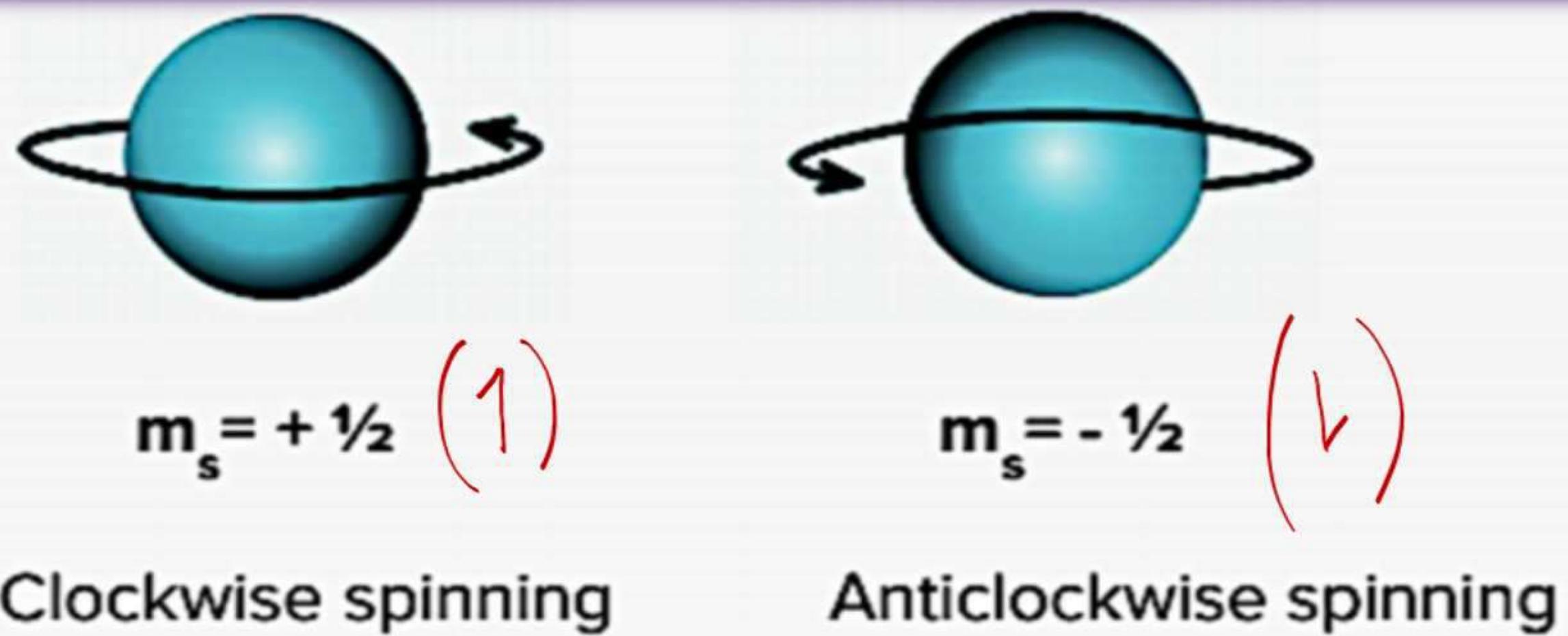


(Degenerate orbital)  $\Rightarrow$  same energy



## Spin quantum numbers(n or m<sub>s</sub>) → from experiment

1. Represents the spin(rotation) of electron around nucleus. It can be CW or ACW
2. It takes value  $\pm \frac{1}{2}$ ,



$$(s) \boxed{1V} = 2e$$

$$(p) \boxed{1V} \boxed{1V} \boxed{1V} = 6e$$

$$(d) \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} = 10e$$

$$(f) \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} \boxed{1V} = 14e$$

## Principal quantum numbers(n)

1. It represents the shell number
2. It determines the size and energy of orbital
3. It tells number of subshells(n) present in it
4. It tells number of orbitals( $n^2$ ) present in it.
5. It tells number of electron  $2(n^2)$  present

## Magnetic quantum numbers(l)

1. Defines the orientation of orbitals in space
2. Its value is from {-l to +l}
3. It tells number of orbitals present in a subshells =  $(2l+1)$
4. Orbitals having same (n,l) value have same energy called **degenerate orbitals**

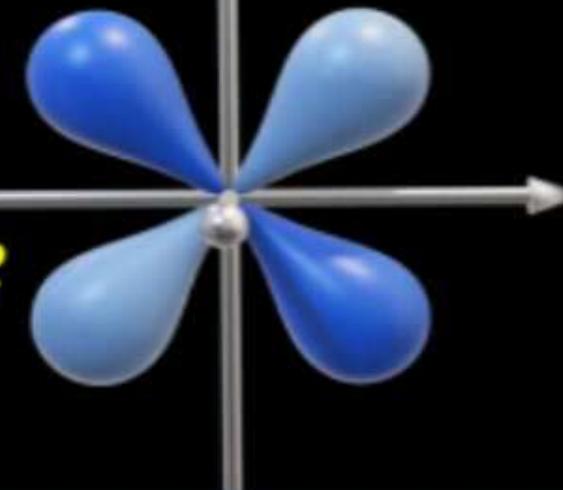
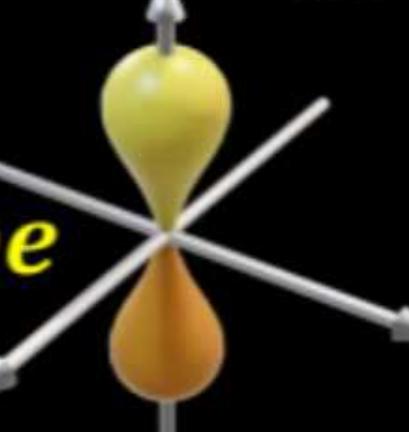
## Spin quantum numbers(s)

1. Represents the spin(rotation) of electron around nucleus. It can be CW or ACW
2. It takes value  $\pm \frac{1}{2}$ ,

## Azimuthal quantum numbers(l)

1. Tells the name, shape and no. of subshell present in given shell
2. Its value is from { l=0 to l=(n-1)}
  - l = 0 → [S] subshell
  - l = 1 → [p] subshell
  - l = 2 → [d] subshell
  - l = 3 → [f] subshell
3. No. of subshell in a shell = (n)
4. **Shape of subshell**
  - ❖ [S] subshell → **Spherical shape**
  - ❖ [p] subshell → **dumbell shape**
  - ❖ [d] subshell → **double dumbell shape**
  - ❖ [f] subshell → **Not in syllabus**

Name of subshells



## Graph predict nodes and antinodes

JEE

- ✓ Nucleus is considered as origin
- ✓ For S-orbital graph starts from max.
- ✓ For p, d, f graph start from origin
- ✓ Point where graph cuts x - axis is called radial node
- ✓ As  $(r)$  increases height of peak decreases

Graph based Schrodinger's Eqn

$$\begin{array}{ll} \textcircled{1} \quad R \text{ v/s } r & \text{or} \quad \psi \text{ v/s } r \\ \textcircled{2} \quad R^2 \text{ v/s } r & \text{or} \quad \psi^2 \text{ v/s } r \end{array}$$

③ Probability ( $P$ ) v/s ( $r$ )

$$\rightarrow (4\pi r^2) R^2 \text{ v/s } (r)$$

Nodes where ( $e^-$ ) finds Prob. is zero

1. The region of zero probability density (excluding nucleus and infinity)
2. Total number of angular nodes(nodal planes) =  $l$
3. Total number of radial nodes =  $n - l - 1$
4. Thus, total number of nodes =  $n - 1$

# Nodes

```

graph LR
    A["# Nodes"] --> B["Radial nodes = (n-1) - l"]
    A --> C["Angular nodes = (l)"]
    A --> D["Total nodes (n-1)"]
  
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Radial nodes =  $[(n-1) - l]$

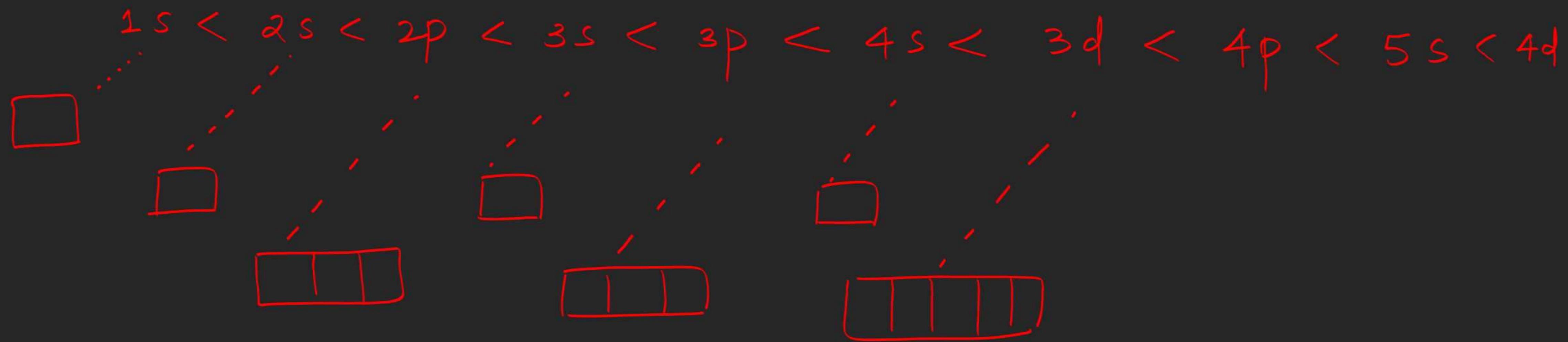
Angular nodes =  $(l)$

Total nodes  $(n-1)$

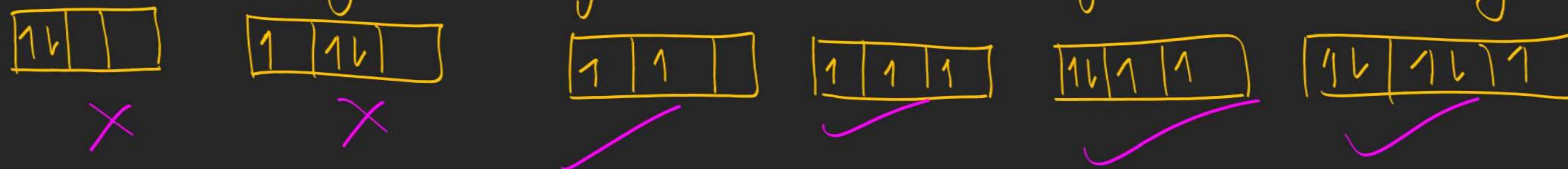
$r$  = distance from nucleus

Rule (1) Aufbau's  $(n+l)$  rule  $\Rightarrow$

- Smaller  $(n+l) \Rightarrow$  smaller energy filled first
- $2f$   $(n+l)$  same }  
they  $(n)$  small }



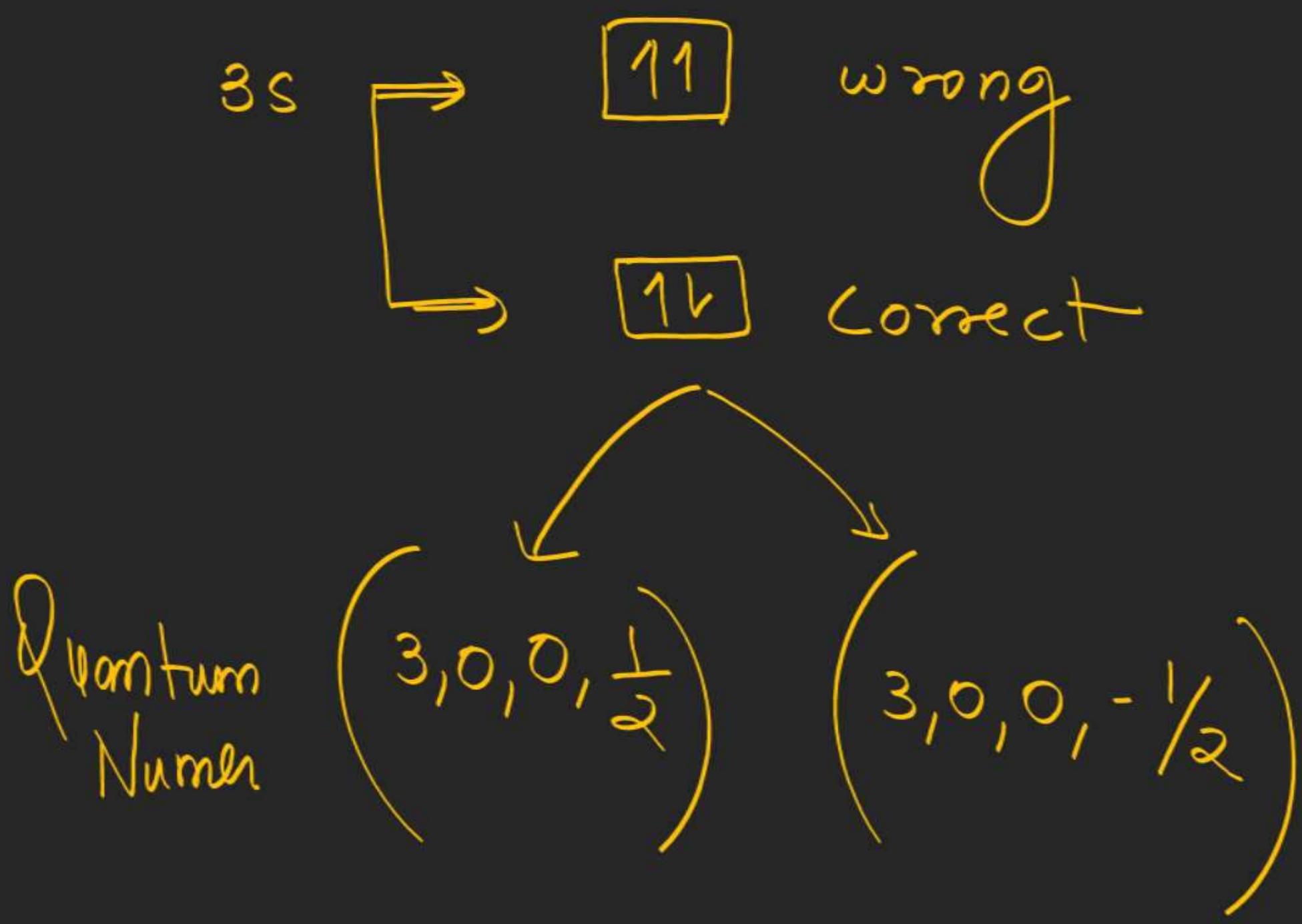
Hund's Rule : Pairing occurs only after each orbital of some subshell gets one e-



### 3) Pauli's exclusion Principle

No two e<sup>-</sup>s can have same set of Quantum no.

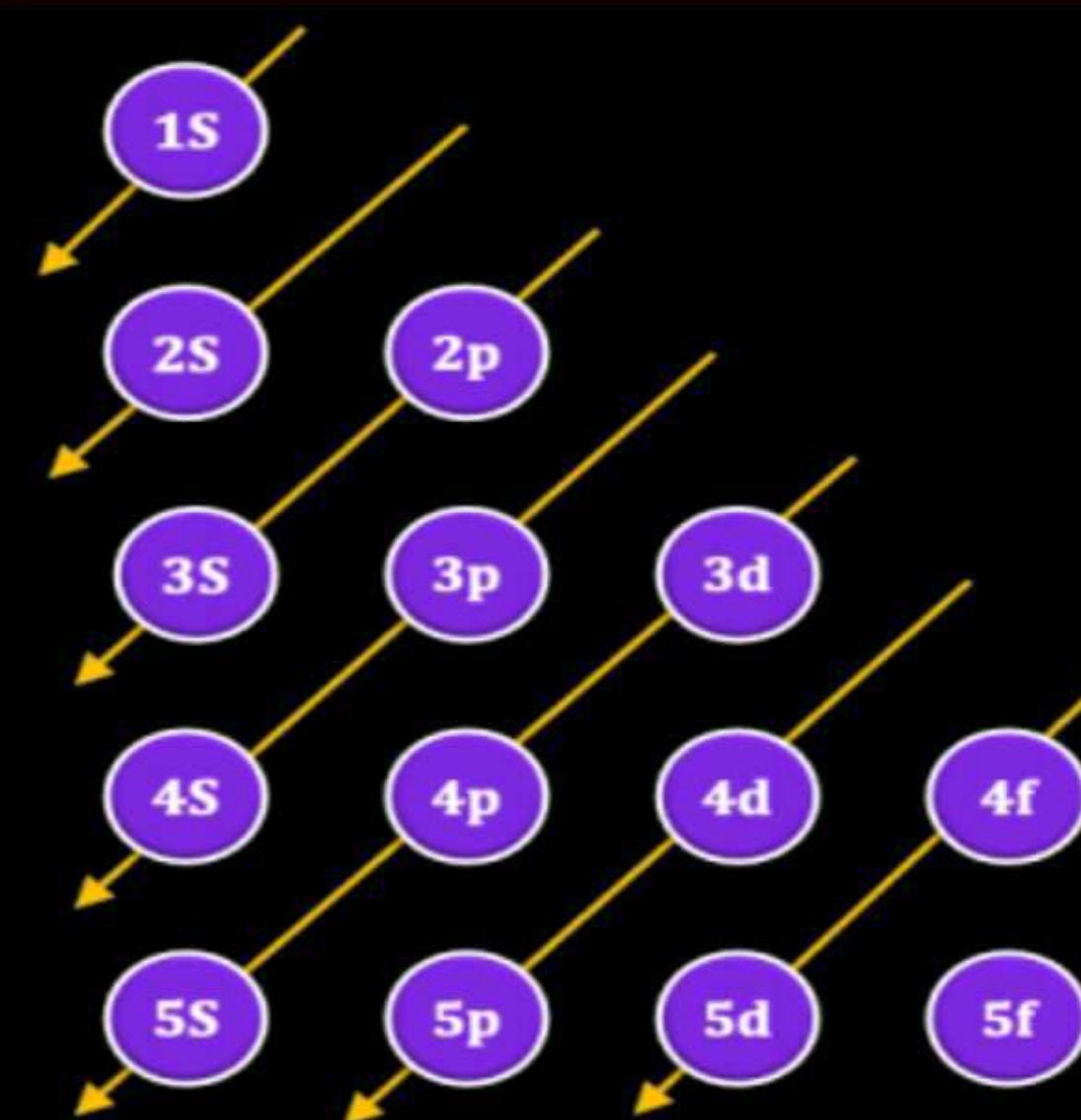
b/c they have diff. spin



# Paired e <sup>-</sup>	# unpaired e <sup>-</sup>	
1L 1L	1 1 1	unpaired e <sup>-</sup> = 3
Paired e <sup>-</sup> = 4		
unpaired = (3)		
(i) Total spin = $n \left( \pm \frac{1}{2} \right)$		Total spin = $\pm 3/2$
* * * (ii) Magnetic Moment = $\sqrt{n(n+2)}$ (Bohr Magneton)		$B \cdot M \cdot = \sqrt{3 \times 5}$ = $\sqrt{15}$ = (3.87)
$n \rightarrow$ no. of unpaired e <sup>-</sup>		

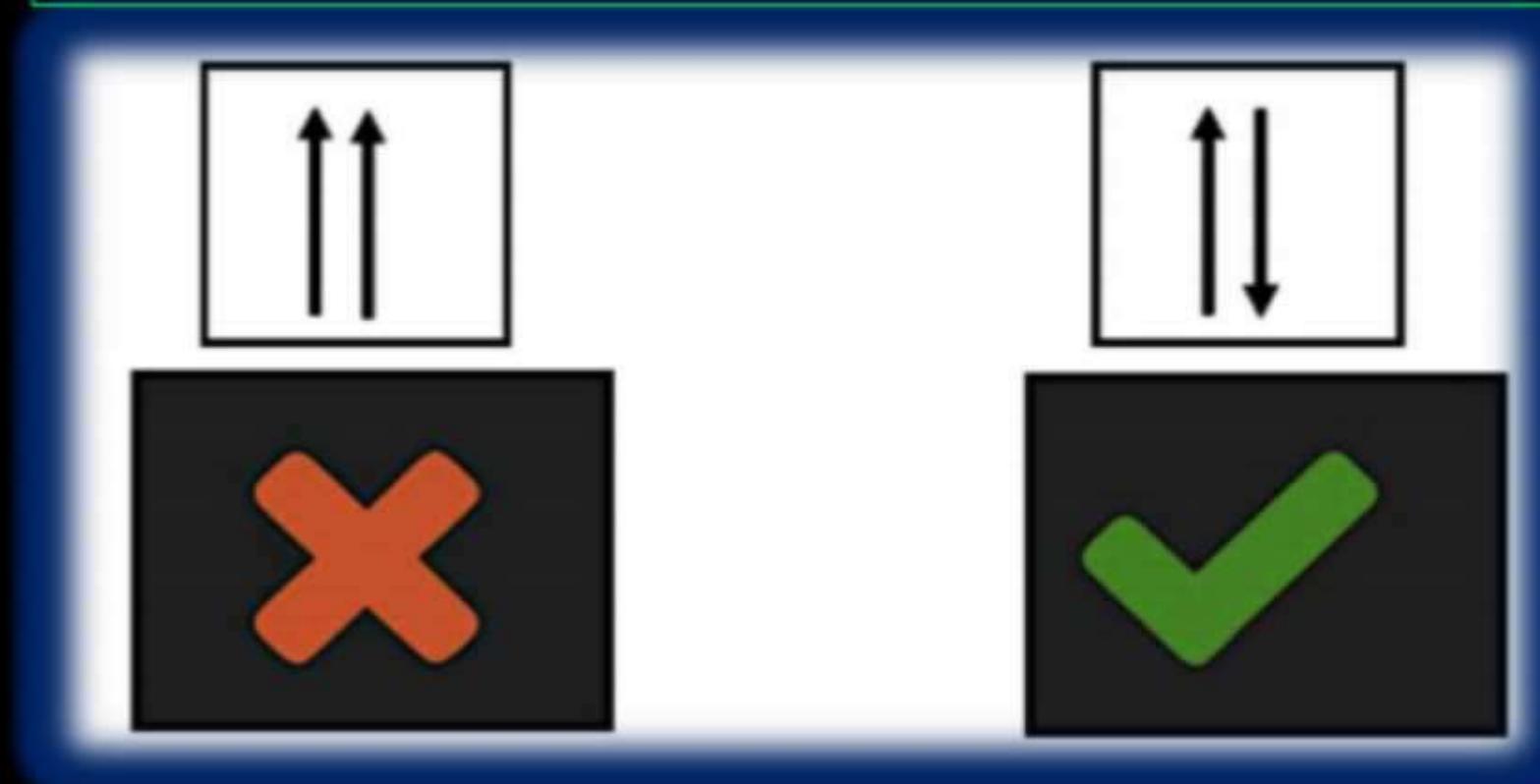
## Aufbau principle

- ❖ Orbitals are filled in order of their increasing energies
- ❖ Orbital with lower value of  $(n+l)$  is filled first
- ❖ If  $(n+l)$  value is same the one having smaller  $(n)$  is filled first.



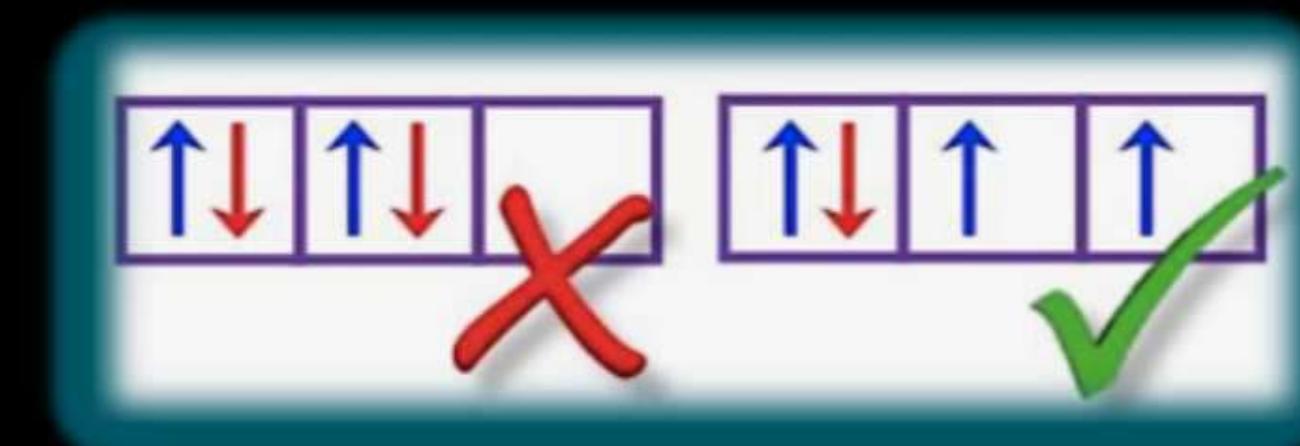
## Pauli exclusion principle

- ❖ No two electrons in an atom can have the same set of four quantum numbers.
- ❖ “Only two electrons may exist in the same orbital and these electrons must have opposite spin.”



## Hund's Rule

- ❖ It states : “pairing of electrons in the orbitals belonging to the same subshell (p, d or f) does not take place until each orbital belonging to that subshell has got one electron each i.e., it is singly occupied”.



## Application of Quantum no.

(i) Comparing energy of orbital

(ii) Electronic configuration

- Para / diamagnetic
- Unpaired  $e^-$  and Magnetic moment =  $\sqrt{n(n+2)}$
- stability of half / fully filled
- excitation of  $e^-$

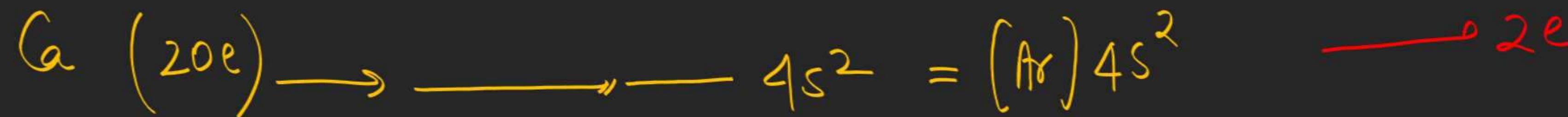
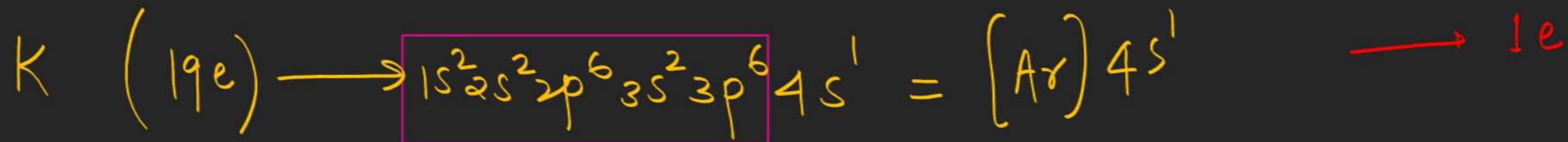
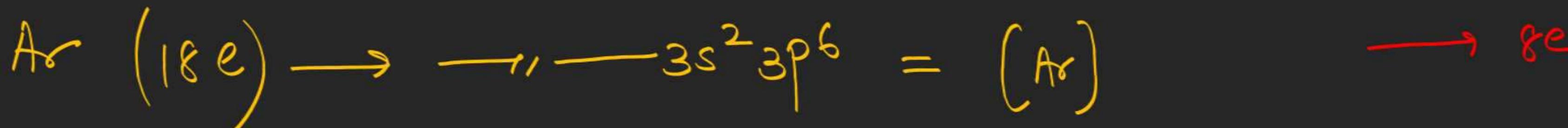
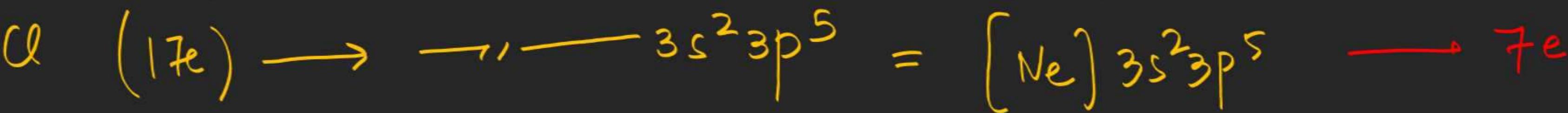
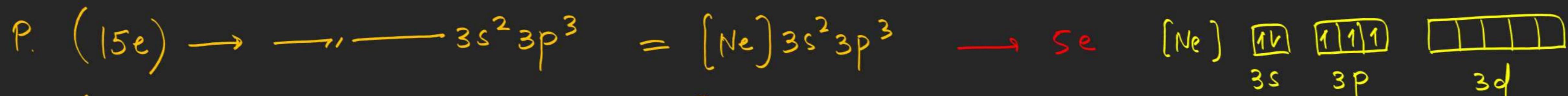
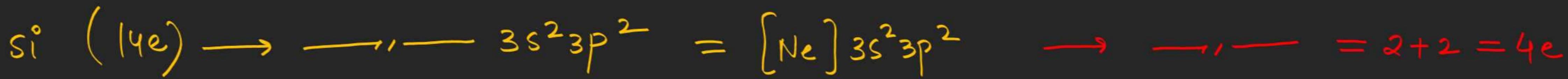
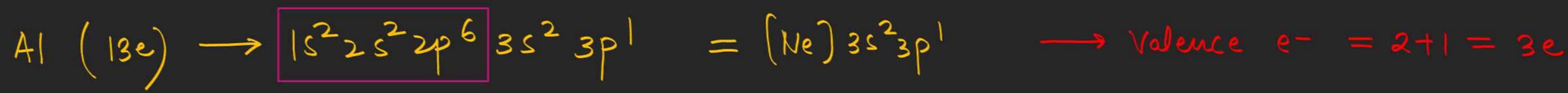
Electronic configuration : Sequential distribution of e- in orbitals using all three rules.

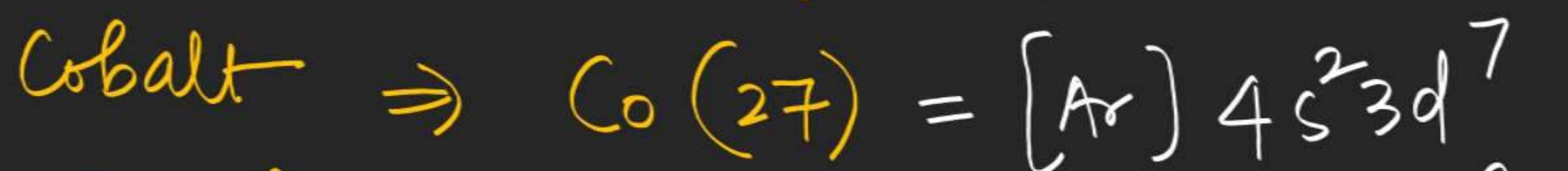
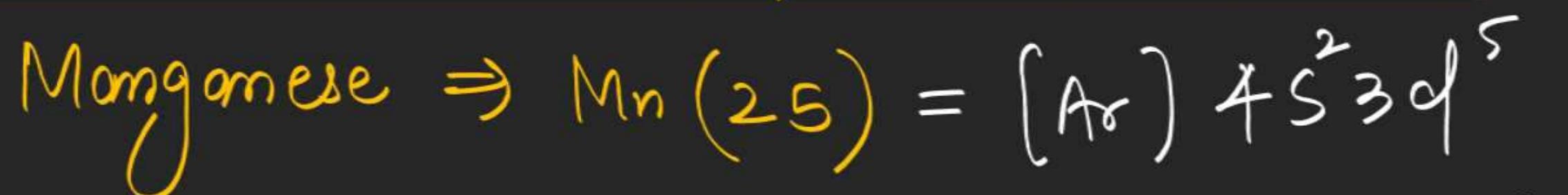
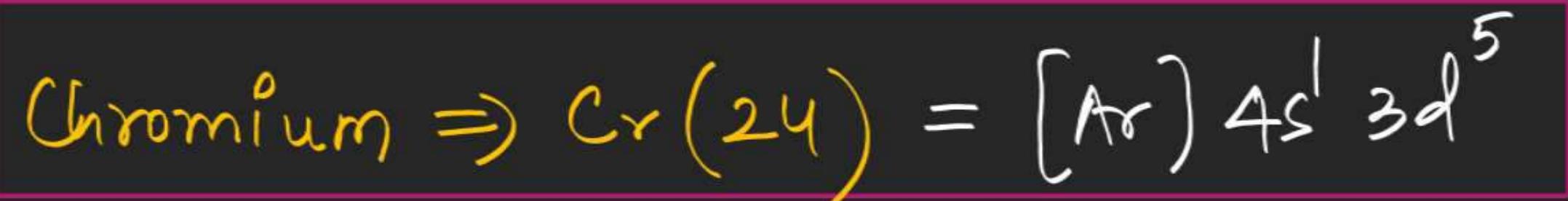
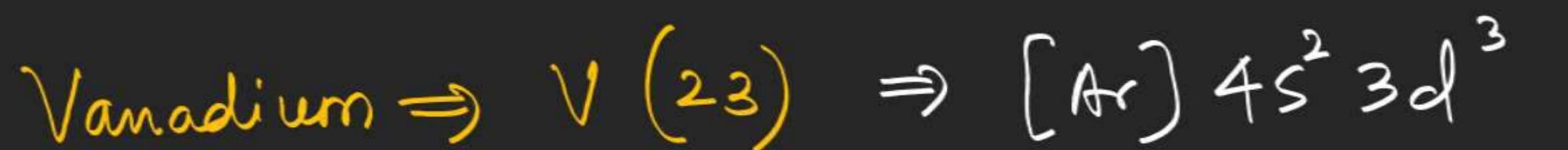
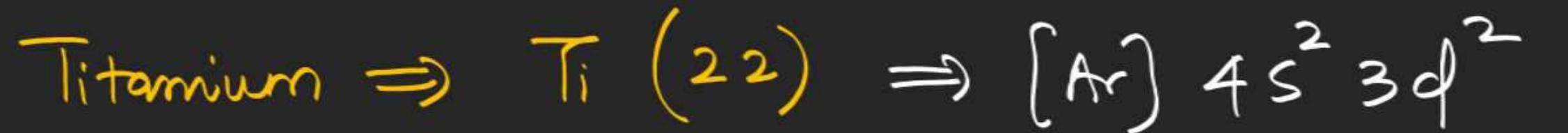
<u>Element</u>	<u>No. of e-</u>	<u>configuration</u>	
H (1e)		$1s^1$ 	1s 2s 2p 3s 3p 4s 3d 4p 5s 4d
He (2e)		$1s^2$ 	
Li (3e)		$1s^2 2s^1$ 	
Be (4e)		$1s^2 2s^2$ 	
B (5e)		$1s^2 2s^2 2p^1$ 	

<u>Element</u>	<u>No. of e-</u>	<u>configuration</u>							
C (6e)	$1s^2 2s^2 2p^2$	<table border="1"><tr><td>1L</td><td>1L</td><td>1 1 1</td></tr></table>	1L	1L	1 1 1				
1L	1L	1 1 1							
N (7e)	$1s^2 2s^2 2p^3$	<table border="1"><tr><td>1L</td><td>1L</td><td>1 1 1</td></tr></table>	1L	1L	1 1 1				
1L	1L	1 1 1							
O (8e)	$1s^2 2s^2 2p^4$	<table border="1"><tr><td>1L</td><td>1L</td><td>1L 1 1</td></tr></table>	1L	1L	1L 1 1				
1L	1L	1L 1 1							
F (9e)	$1s^2 2s^2 2p^5$	<table border="1"><tr><td>1L</td><td>1L</td><td>1L 1L 1</td></tr></table>	1L	1L	1L 1L 1				
1L	1L	1L 1L 1							
Ne (10e)	$1s^2 2s^2 2p^6$	<table border="1"><tr><td>1L 1s</td><td>1L 2s</td><td>1L 1L 1L 2p</td></tr></table>	1L 1s	1L 2s	1L 1L 1L 2p				
1L 1s	1L 2s	1L 1L 1L 2p							
Na (11e)	$1s^2 2s^2 2p^6 3s^1$	<table border="1"><tr><td>1L 1s</td><td>1L 2s</td><td>1L 1L 1L 2p</td><td>1 3s</td><td>1 1 1 3p</td><td>1 1 1 1 3d</td></tr></table>	1L 1s	1L 2s	1L 1L 1L 2p	1 3s	1 1 1 3p	1 1 1 1 3d	$4p$ $5s$ $4d$ $1$ $1$
1L 1s	1L 2s	1L 1L 1L 2p	1 3s	1 1 1 3p	1 1 1 1 3d				
Mg (12e)	$1s^2 2s^2 2p^6 3s^2$	<table border="1"><tr><td>1L 1s</td><td>1L 2s</td><td>1L 1L 1L 2p</td><td>1L 3s</td><td>1 1 1 3p</td><td>1 1 1 1 3d</td></tr></table>	1L 1s	1L 2s	1L 1L 1L 2p	1L 3s	1 1 1 3p	1 1 1 1 3d	
1L 1s	1L 2s	1L 1L 1L 2p	1L 3s	1 1 1 3p	1 1 1 1 3d				

Noble Gas

$He \rightarrow 2e$	1s
$Ne \rightarrow 10e$	2s
$Ar \rightarrow 18e$	2p
$Kr \rightarrow 36e$	3s
$Xe \rightarrow 54$	3p
	4s
	3d
	4p
	5s
	4d
	1
	1





# Half / fully filled  
are more stable  
orbitals

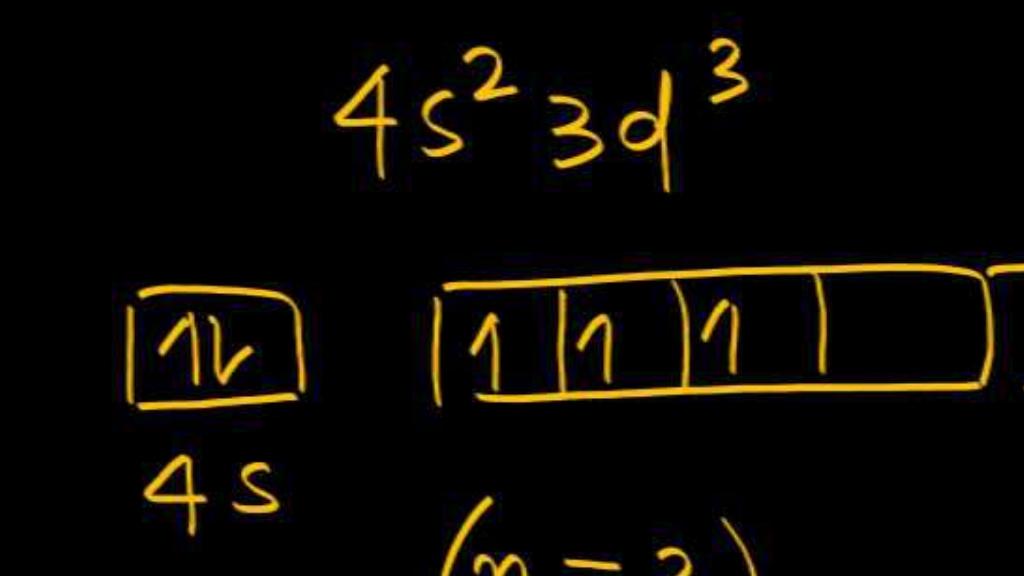
d<sup>5</sup>, d<sup>10</sup>

p<sup>6</sup>, p<sup>3</sup>

are stable  
configuration

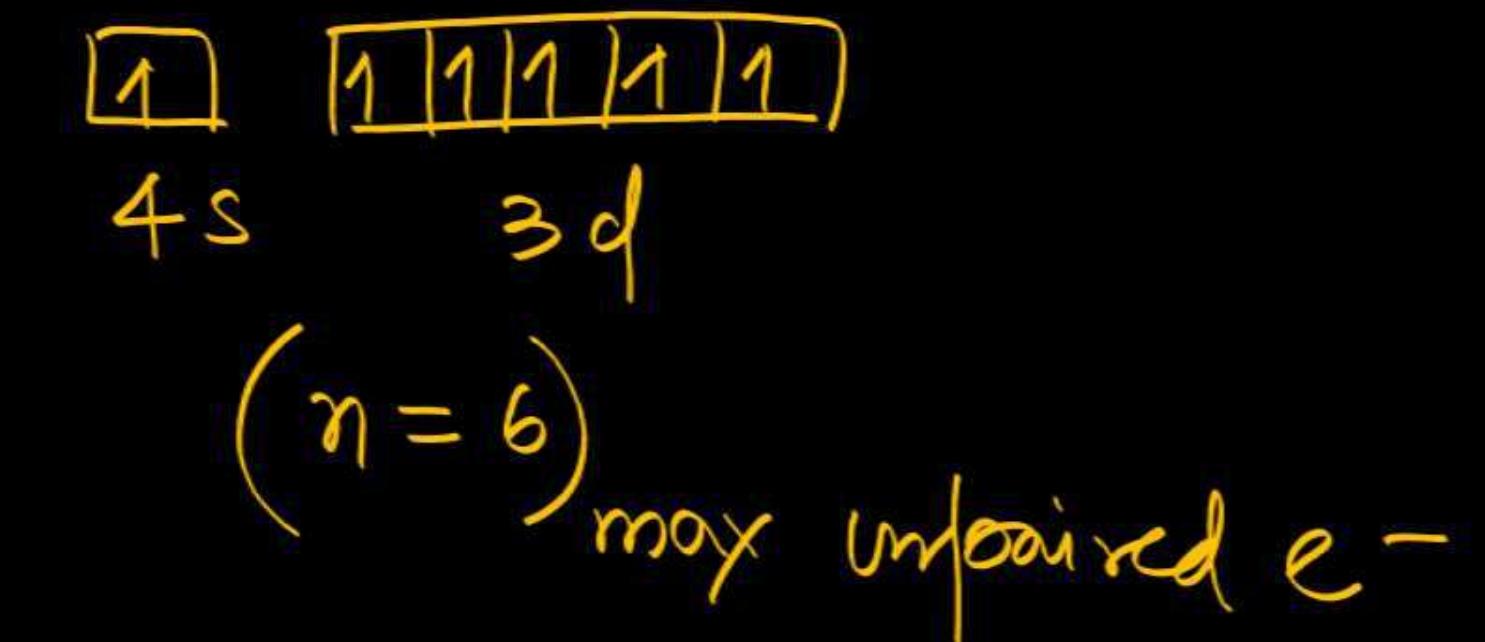
**Sc (21)****Ti****V (23)****Cr<sup>+</sup> (24)****Mn**

$$4s^2 3d^1$$



→ unpaired e-

$$4s^1 3d^5$$

**Fe (26)****Co****Ni (28)****Cu****Zn**

$$4s^2 3d^6$$



$n=3$

$$4s^2 3d^8$$

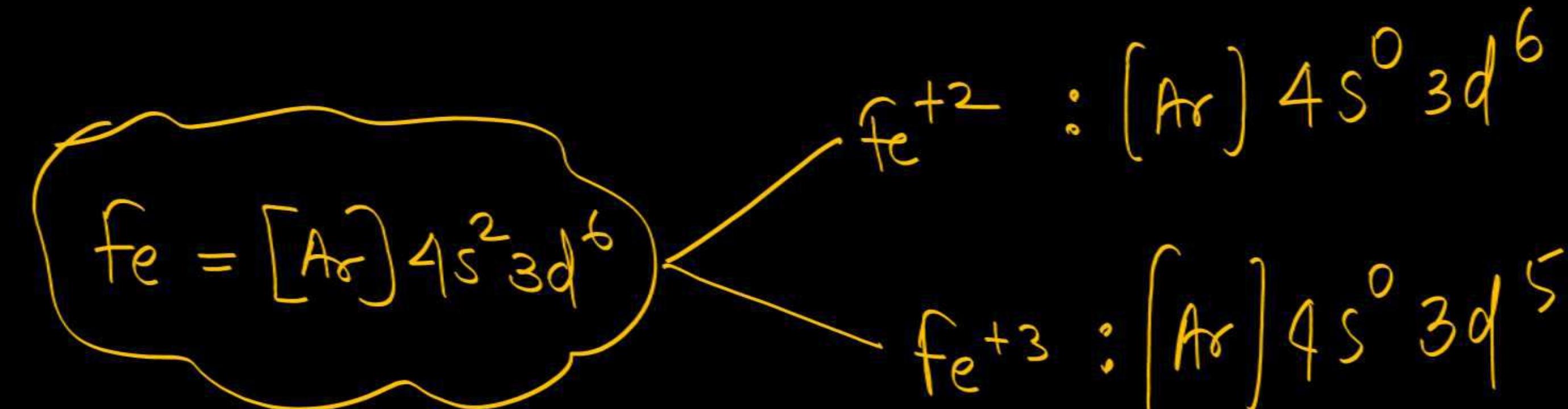
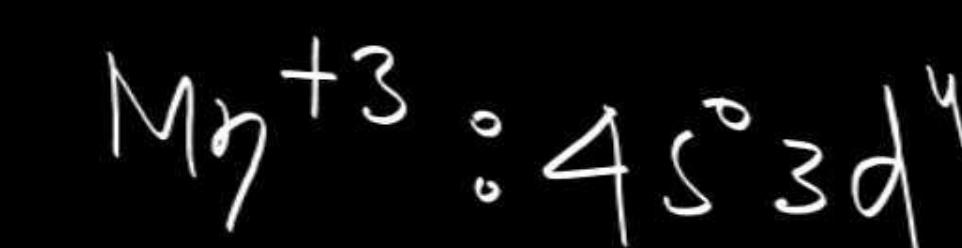
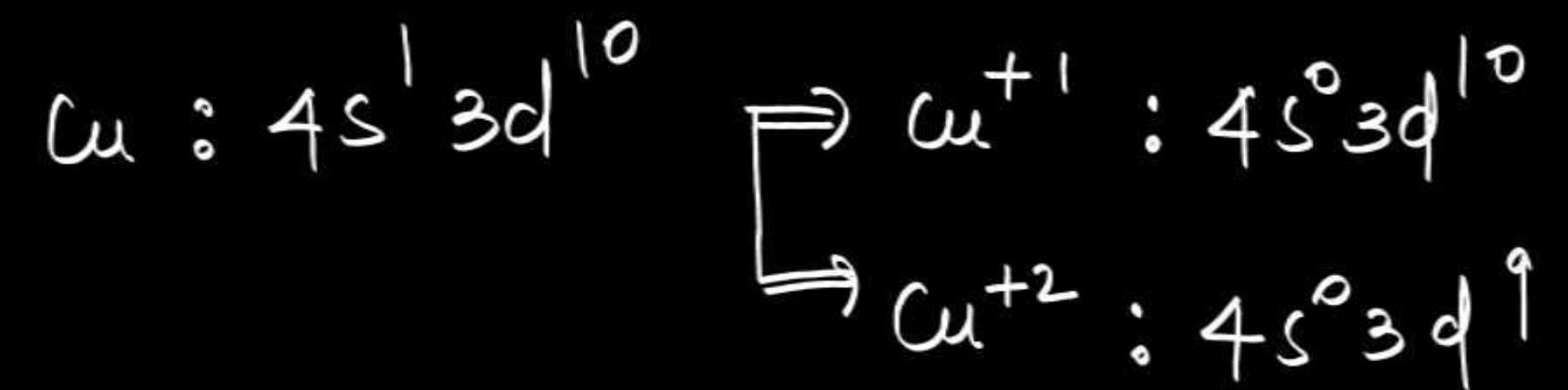
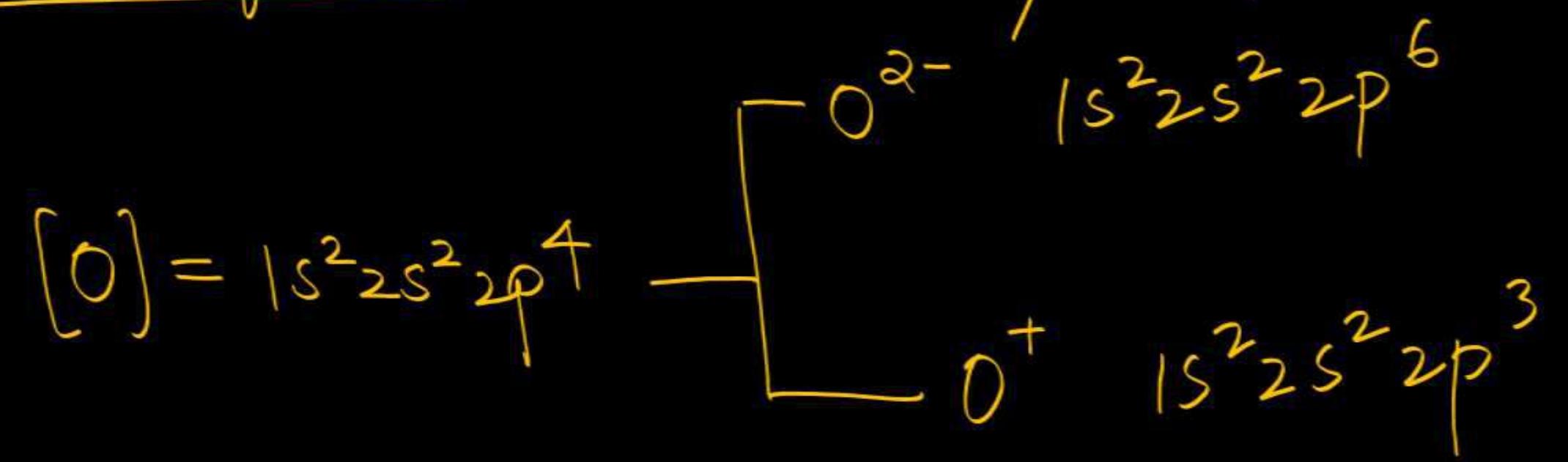


$n=1$

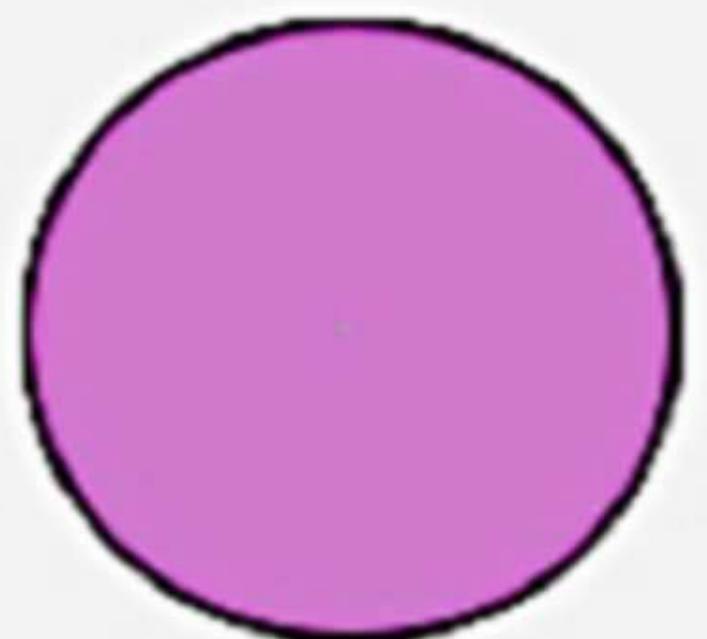
$n=0$

Sc	Sc <sup>+</sup>	Sc <sup>+2</sup>	V (23)	V <sup>+</sup>	V <sup>+3</sup>
4s <sup>2</sup> 3d <sup>1</sup>	4s <sup>1</sup> 3d <sup>1</sup>	4s <sup>0</sup> 3d <sup>1</sup>	4s <sup>2</sup> 3d <sup>3</sup>	4s <sup>1</sup> 3d <sup>3</sup>	4s <sup>0</sup> 3d <sup>2</sup>

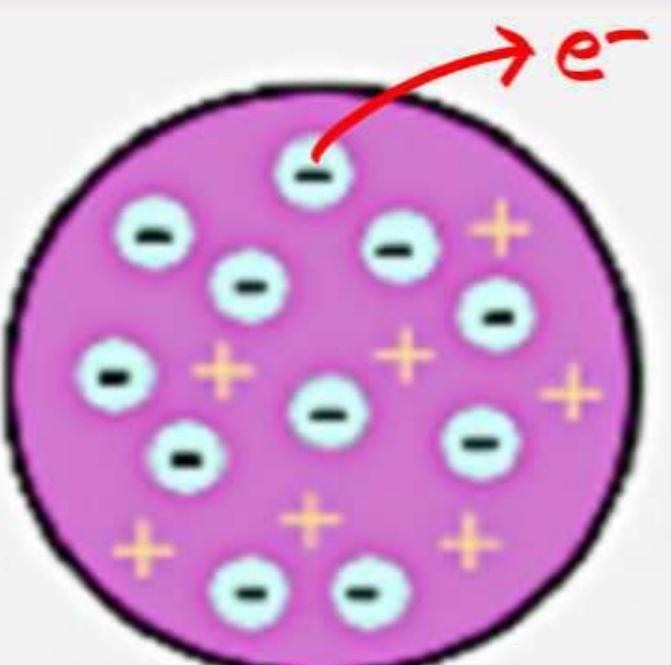
configuration of ions: e<sup>-</sup> are removed / added in outer shell



# Atomic Models We know



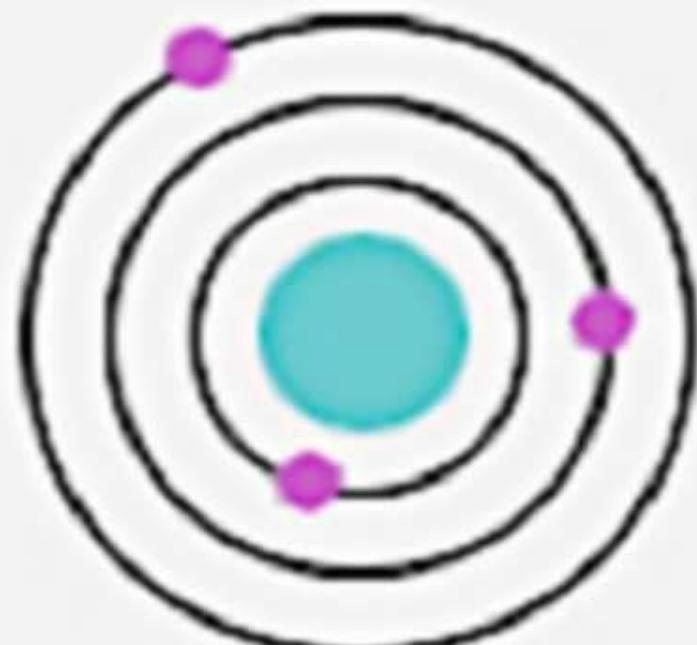
Dalton  
"Billiard Ball" Model



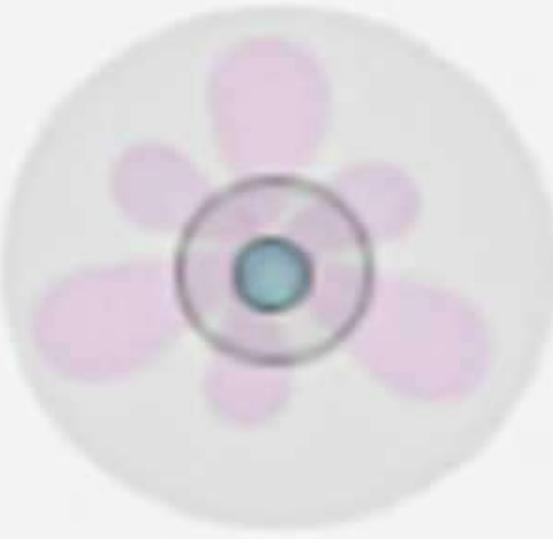
Thomson  
"Plum Pudding" Model



Rutherford Model



Bohr Model



Quantum Mechanical  
Model

# Smallest particle (Atom)

# No subatomic particle

# e<sup>-</sup> at rest

$$\xrightarrow{\quad} \mathbf{f}_{\text{net}} = 0$$

\* Equilibrium

# Discovered nucleus

#  $\alpha$ - gold foil exp.

# e<sup>-</sup> revolves  
around nucleus

# ( $n, p$ ) inside nucleus

Classical  
Mechanics



NLM  
applicable

(Macroscopic  
object)

Quantum  
Mechanics

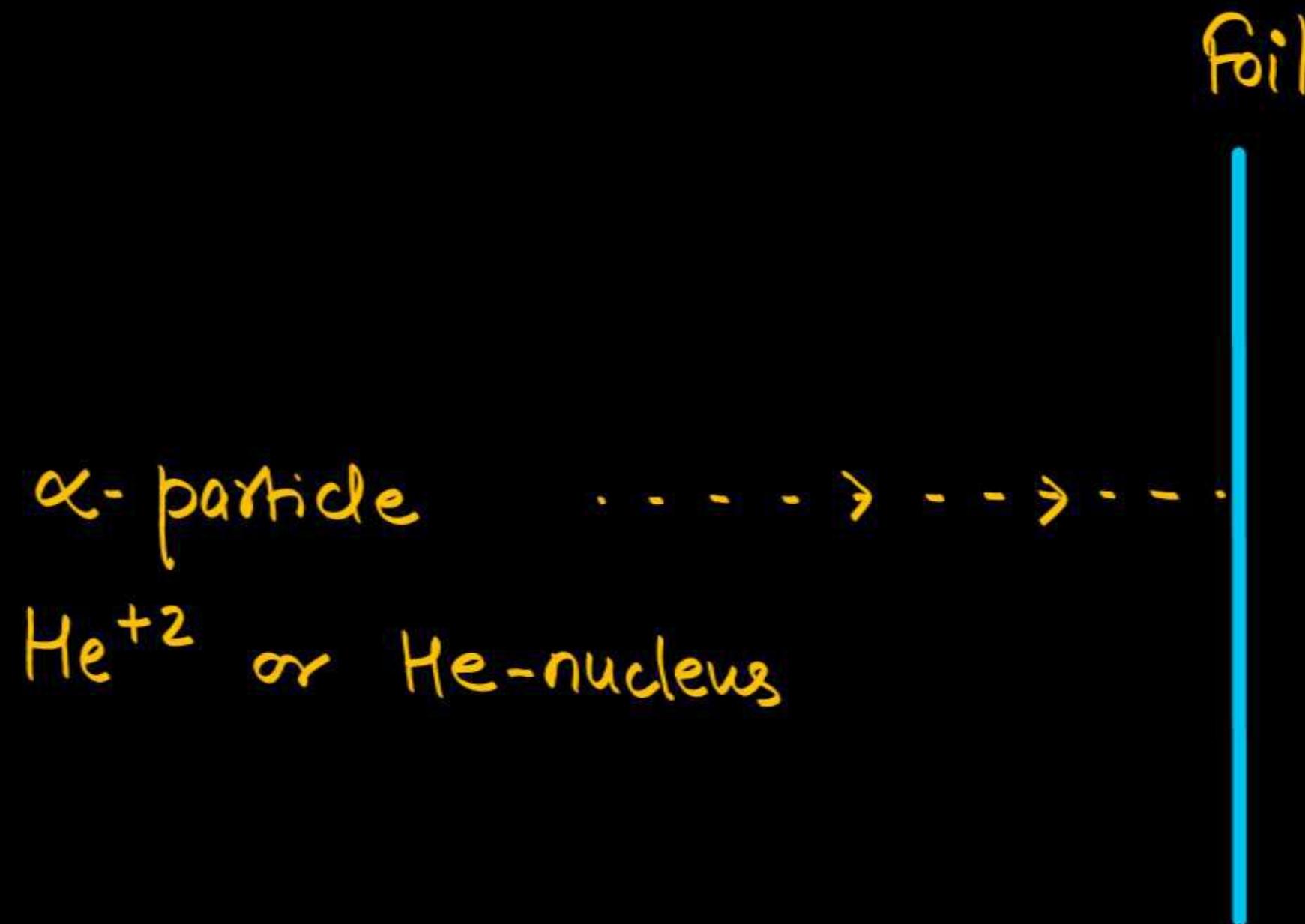


NLM  
not applicable

(microscopic  
object)

# Rutherford Model

❖ No. of  $\alpha$  - particle deflected  $\propto$  thickness of foil  $\propto$  (atomic number) $^2$   $\propto \frac{1}{(K.E. \text{ of } \alpha\text{-particle})^2} \propto \frac{1}{[\sin \frac{\theta}{2}]^4}$



$$(i) \frac{N_1}{N_2} = \frac{t_1}{t_2} \quad \left( t = \text{thickness} \right)$$

$$(ii) \frac{N_1}{N_2} = \left( \frac{Z_1}{Z_2} \right)^2$$

$$(iii) \frac{N_1}{N_2} = \left( \frac{KE_2}{KE_1} \right)^2$$

$$(iv) \frac{N_1}{N_2} = \left( \frac{\sin \theta_2/2}{\sin \theta_1/2} \right)^4$$

Ex : If 50  $\alpha$ -particle were deflected using mg- foil of thickness 4mm.  
then no. of Particles deflected

a) thickness is doubled

b) some thickness of 'cr'

c) Double  $(KE)$

$$\rightarrow \frac{N_1}{N_2} = \left( \frac{KE_2}{KE_1} \right)^2$$

$$\frac{50}{N_2} = \left( \frac{2}{1} \right)^2$$

$$N_2 = 50/4$$

$$a) \frac{N_1}{N_2} = \frac{t_1}{t_2} \Rightarrow \frac{50}{N_2} = \frac{4}{8}$$

$$N_2 = 100$$

$$b) \frac{N_1}{N_2} = \left( \frac{Z_1}{Z_2} \right)^2 \Rightarrow \frac{50}{N_2} = \left( \frac{12}{24} \right)^2$$

$$200 = N_2$$

Wave: Transfer of energy / momentum / Pressure difference

G.K.

Type of waves

LIVE

Mechanical Waves

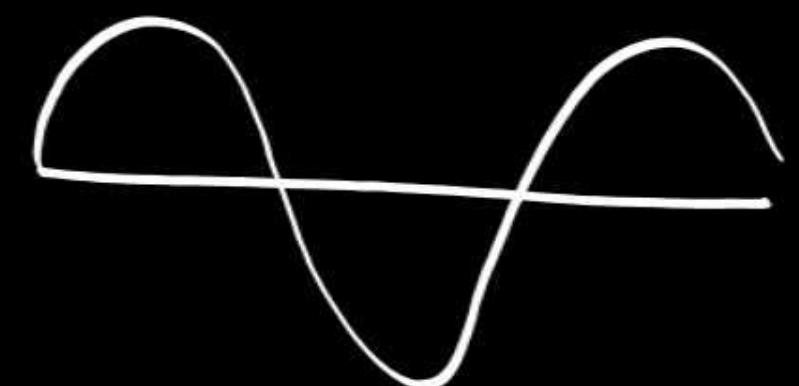
(EMW)

Electromagnetic Waves

(Non-mechanical wave)

\* Require medium to travel

Transverse



ex String wave

Longitudinal

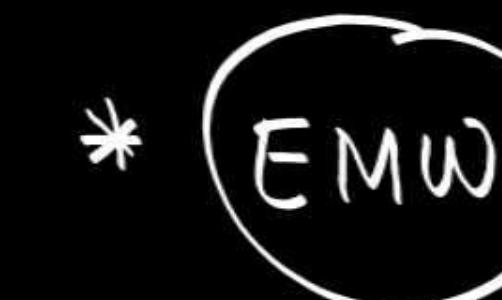


Sound wave

\* Doesn't req. medium to travel

\* Transverse wave

\* Produced by accelerating charge particle



\* EMW is an oscillation of electric ( $\vec{E}$ ) and magnetic field vectors  $\vec{B}$  to its propagation

## Wave Parameter

Speed of All  
EM waves is

Speed of  
light ( $c$ ) =  $3 \times 10^8 \text{ m/s}$

(i) wavelength ( $\lambda$ )

(ii) wave no. ( $\frac{1}{\lambda}$ )

(iii) Time period ( $T = \frac{\lambda}{c}$ )

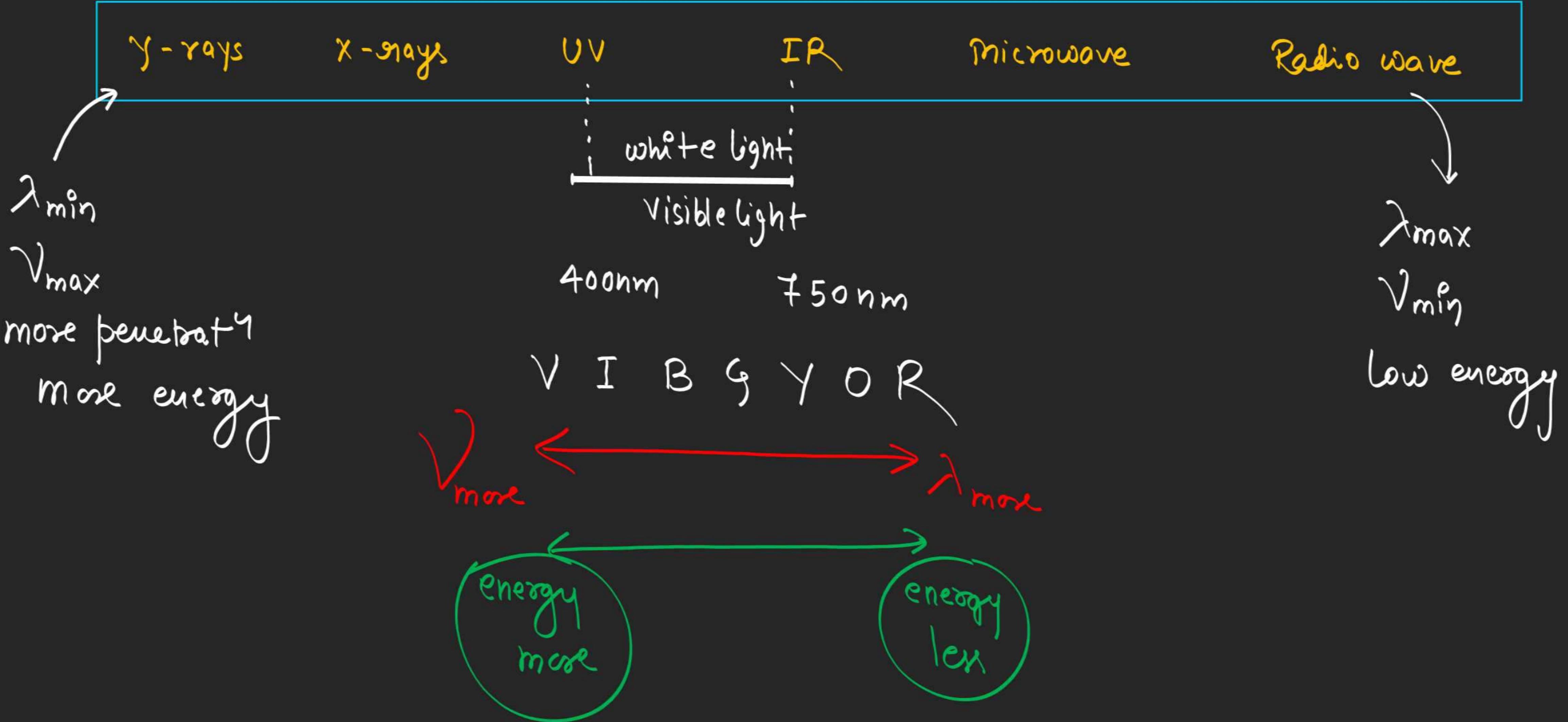
(iv) frequency ( $\nu$ ) =  $\frac{1}{T}$

↳ Unit : Hz

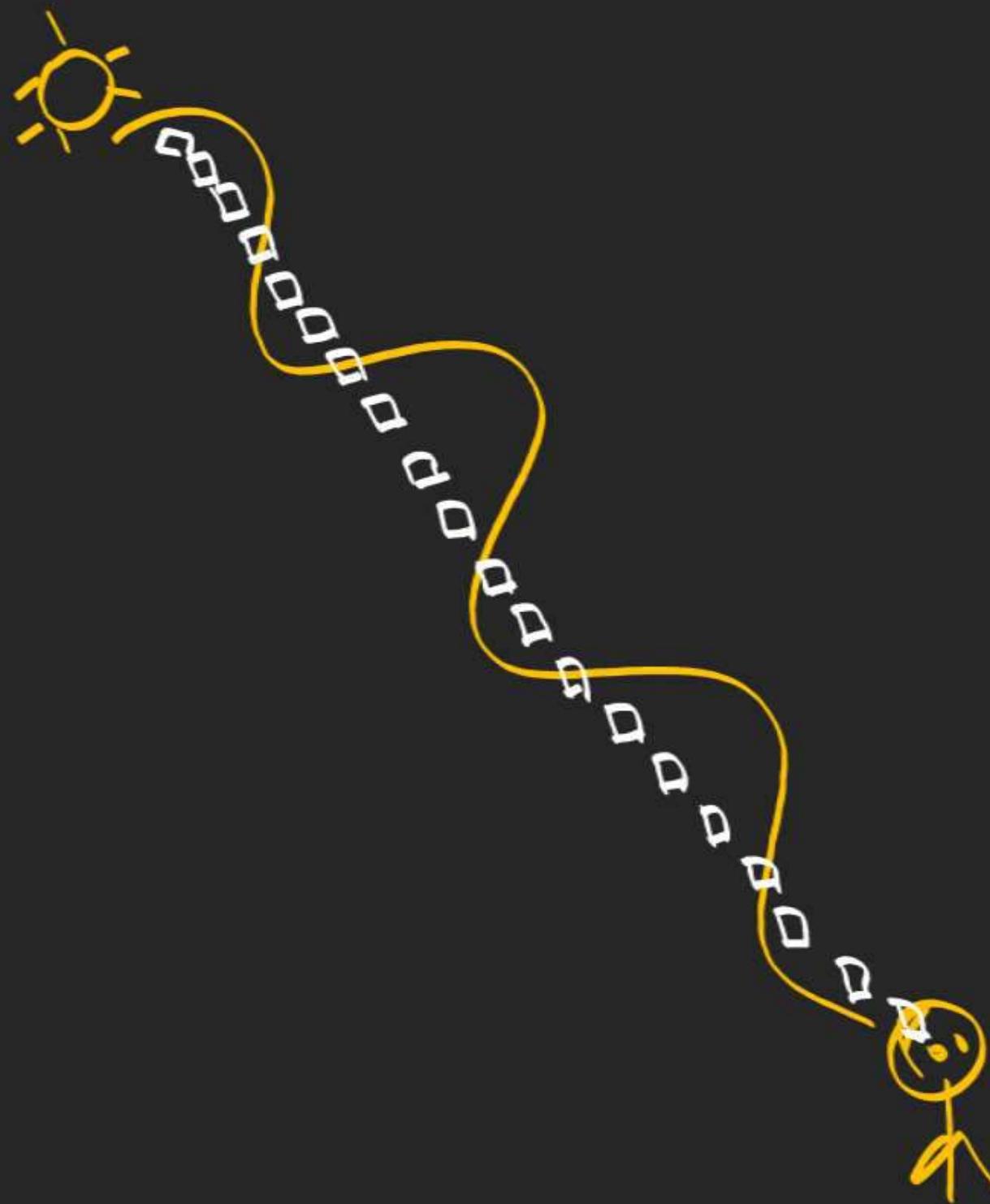
  $\nu = \frac{c}{\lambda}$

$$\left. \begin{array}{l} * \text{ nm} = 10^{-9} \text{ m} \\ * \text{ Å} = 10^{-10} \text{ m} \\ * \text{ pm} = 10^{-12} \text{ m} \end{array} \right\}$$

Wave Number ( $\bar{\nu}$ ) : It is defined as the number of wavelengths per unit length.



Planck's Quantum Theory : energy transfer through rad<sup>n</sup> is not continuous but in **discrete manner**



→ in the form of small packets of energy called photon (Quanta of light)

→ Quantisation of energy  
(energy is integral multiple of one photon energy)

Energy of photon  $\propto$  freq  $\propto \frac{1}{\text{Wavelength}}$

$$E = h\nu = \frac{hc}{\lambda}$$

where  
 $h$  = Planck's const.

$$= 6.62 \times 10^{-34} \text{ J.s.}$$

$$\Rightarrow \text{Total energy} \Rightarrow E_T = n(h\nu) = n \frac{hc}{\lambda}$$

( $n \rightarrow$  no. of photon)

# Equating energy of photons and wave

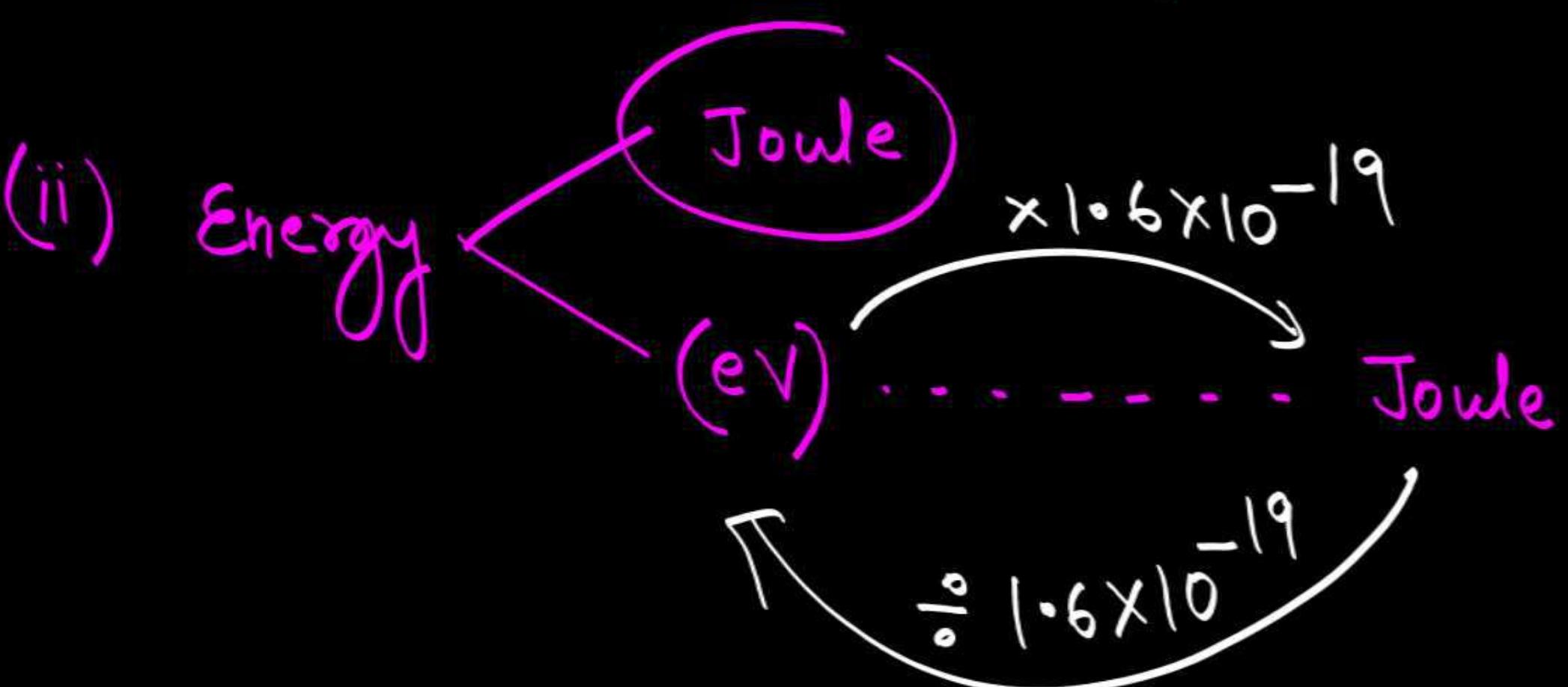
(i) one photon :  $\epsilon = h\nu = \frac{hc}{\lambda}$

(ii) n-photon :  $\epsilon_T = n(h\nu) = n\left(\frac{hc}{\lambda}\right)$

(iii) Power in 1 sec = Energy =  $n(h\nu) = n\left(\frac{hc}{\lambda}\right)$

(iv)  $(\text{Power} \times \text{time}) = \frac{\text{energy}}{\text{total}}$  =  $\dots$

i) 1 mole photon =  $6.02 \times 10^{23}$  photon  
 $= (N_A)$  photon  
 $\Rightarrow \text{energy} = (N_A) h\nu$   
 $= (N_A) \frac{hc}{\lambda}$

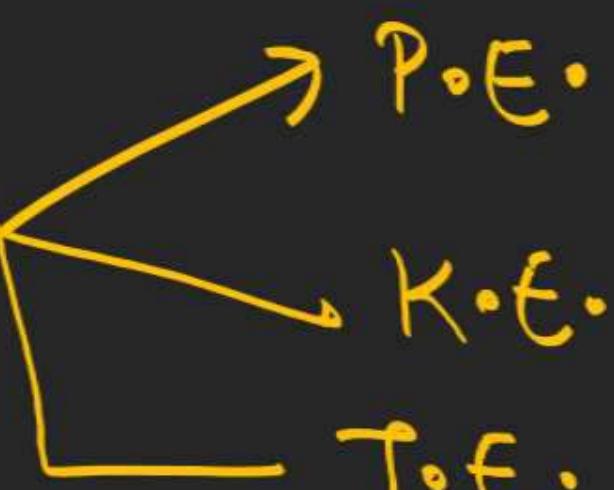


Bohr model (Valid for single e- system)  $\longrightarrow$  H, He<sup>+</sup>, Li<sup>+2</sup>, Be<sup>+3</sup> . . .

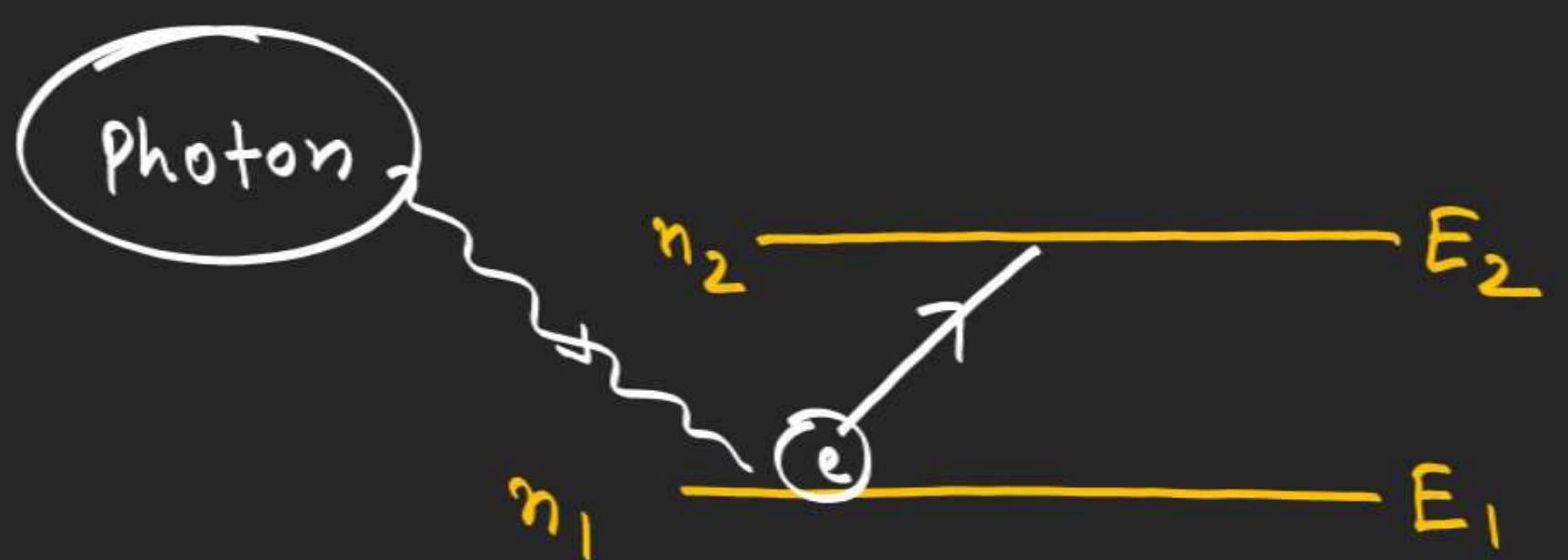
(i) e- will revolve only in those orbit which has angular momentum =  $n \left( \frac{h}{2\pi} \right)$

$$\Rightarrow mvr = \frac{nh}{2\pi} \quad \text{where } n=1, 2, 3, 4, \dots$$

$\Rightarrow$  Quantisation of ang. mom.

(ii) stationary orbits have fixed energy of e-   
fixed velocity of e-  
fixed radius of orbit

\*



④ e- from higher energy level jump to lower energy level emit grad<sup>n</sup> (Photon)

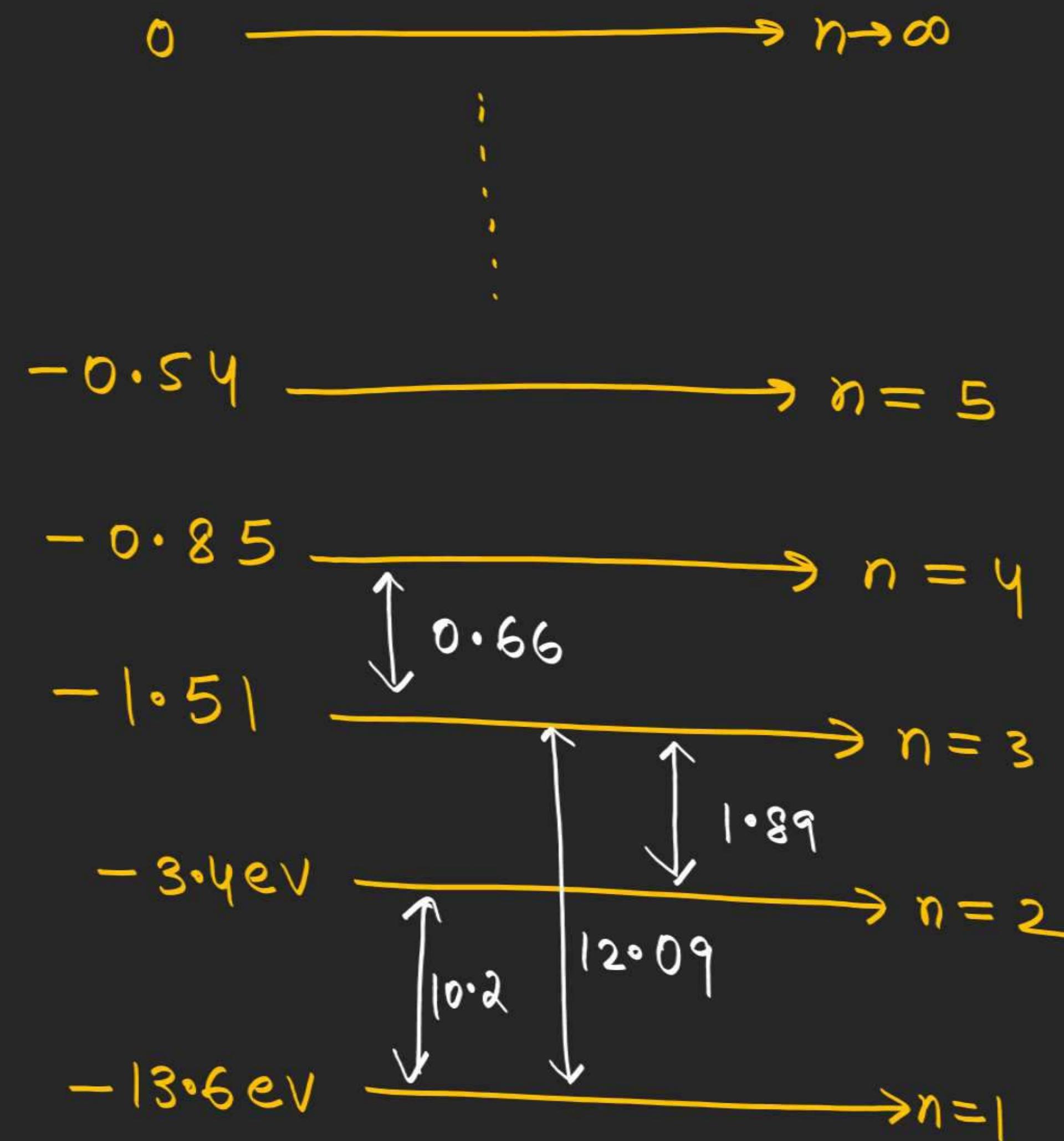
# Photon / rad<sup>n</sup> emitted energy = energy gap

(e-) upon absorbing energy  
Jumps into higher energy level.

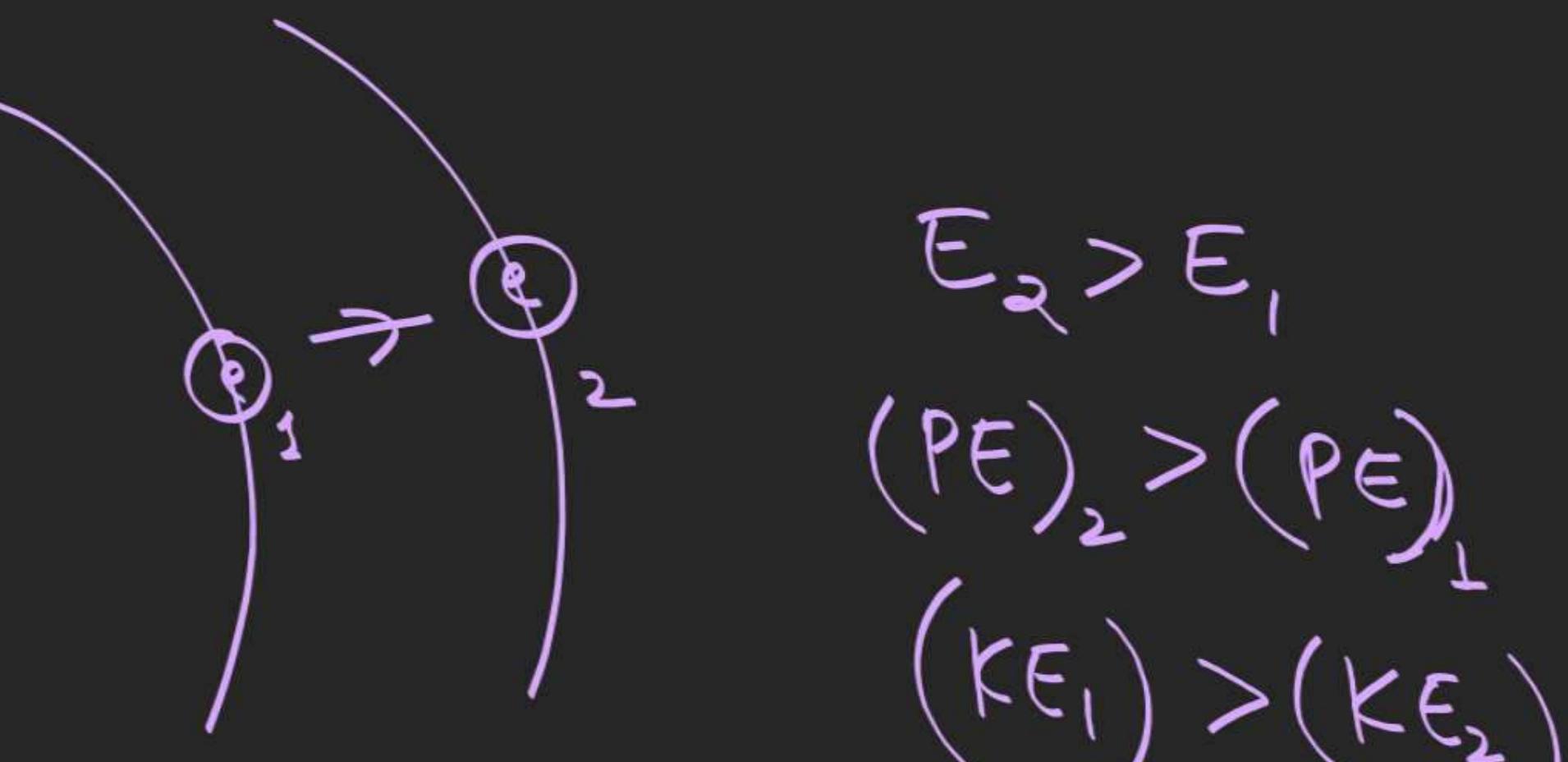
A diagram illustrating the absorption process. Two horizontal lines represent energy levels, with the upper one being higher than the lower one. A small circle with an 'e' inside is shown on the upper level. An arrow points downwards from the upper level to the lower level, labeled  $\Delta E$  between the levels. A wavy line labeled 'Photon' originates from the electron at the upper level and extends to the right. To the right of the photon line, the equation  $\Delta E = h\nu$  is written, followed by  $= \frac{hc}{\lambda}$  and  $= \frac{1240}{\lambda(\text{nm})} \text{ ev}$ .

$$\Delta E = h\nu$$
$$= \frac{hc}{\lambda}$$
$$= \frac{1240}{\lambda(\text{nm})} \text{ ev}$$

$$\text{Energy} = -13.6 \left(\frac{Z}{n}\right)^2$$



# As  $(n) \uparrow \Rightarrow (\text{Energy}, \text{P.E.}) \uparrow$   
 $\Rightarrow (KE) \downarrow$



$$\begin{aligned} E_2 &> E_1 \\ (PE)_2 &> (PE)_1 \\ (KE_1) &> (KE_2) \end{aligned}$$

Note:

5th excited state  $\rightarrow n = 6$

2nd excited state  $\rightarrow n = 3$

1st excited state  $\rightarrow n = 2$

Ground state  $\rightarrow n = 1$

Note

Ionisation energy

Energy required to remove  $e^-$  from  $n = 1$  to  $n \rightarrow \infty$

$$I \cdot E = E_{\infty} - E_1$$

$$= 0 + 13.6 \frac{z^2}{(1)^2}$$

$$\boxed{I \cdot E = 13.6 z^2}$$



Ionisation Potential : Potential diff to remove e- from  $n=1$  to  $n \rightarrow \infty$

+ve

$$I.P. = 13.6 Z^2 \text{ volt}$$

Excitation potential

{ From  $n_1 \rightarrow n_2$

Excitation energy

+ve Value

$$13.6 Z^2 \left[ \frac{1}{n_2^2} - \frac{1}{n_1^2} \right]$$

Separation energy from given ( $n$ ) to  $\infty \Rightarrow \left( 13.6 \frac{Z^2}{n^2} \right)$

$$(i) E_n = -13.6 \frac{Z^2}{n^2}$$

$$(ii) KE = 13.6 \frac{Z^2}{n^2}$$

$$(iii) PE = -27.2 \frac{Z^2}{n^2}$$

$$(iv) I.E. = 13.6 Z^2 \quad \& \quad (I.P.)$$

$$(v) \text{sep'n energy} = 13.6 \frac{Z^2}{n^2} \quad \& \quad (\text{separation potential})$$

$$(vi) \text{Excitation energy} = 13.6 Z^2 \left( \frac{1}{n_2^2} - \frac{1}{n_1^2} \right)$$

$$(vii) \Delta E = \underline{\hspace{2cm}} - \underline{\hspace{2cm}}$$

$$(viii) \frac{E_1}{E_2} = \left( \frac{Z_1}{Z_2} \right)^2 \left( \frac{n_2}{n_1} \right)^2$$

(ix)  $n \uparrow \quad KE \downarrow$   
 $(\text{Energy, PE}) \uparrow$

Radius of orbit :  $r_n = 0.529 \left( \frac{n^2}{Z} \right) A^\circ$

(i) Compare :  $\frac{r_1}{r_2} = \left( \frac{n_1}{n_2} \right)^2 \left( \frac{Z_2}{Z_1} \right)$

(ii) Bohr radius ( $n=1$ )  $\xrightarrow{\text{H-atom}} 0.529 \times \frac{1^2}{1} = 0.529$

(iii) If Bohr radius is ( $x$ )  
then radius of  $n=3$  }  $\left. \begin{array}{l} \\ \end{array} \right\} \text{Ans } (9x)$

find ratio of radius of  $n=3$  for  $\text{He}^+$

and  $4m$  excitate state of  $\text{Be}^{+3}$

Ans.  $\frac{\gamma_1}{\gamma_2} = \left(\frac{n_1}{n_2}\right)^2 \left(\frac{z_2}{z_1}\right)$

$$= \left(\frac{3}{5}\right)^2 \left(\frac{4}{2}\right)$$

$$= \frac{9}{25} \times 2 = \frac{18}{25}$$

Velocity of  $e^-$  in an orbit

$$V_n = 2.18 \times 10^6 \left( \frac{z}{n} \right) \text{ m/s}$$

Time period of  $e^-$

$$\textcircled{*} \quad T = \frac{2\pi r}{v}$$

#  $\frac{V_1}{V_2} = \left( \frac{z_1}{z_2} \times \frac{n_2}{n_1} \right)$

$$T = \frac{2\pi \left( 0.529 \times \frac{n^2}{z} \right) \times 10^{-10}}{2.18 \times 10^6 \left( \frac{z}{n} \right)}$$

omp

$$T \propto \frac{n^3}{z^2} \Rightarrow \frac{T_1}{T_2} = \left( \frac{n_1}{n_2} \right)^3 \left( \frac{z_2}{z_1} \right)^2$$

Angular mom:

$$\vec{L} = mv\vec{\tau} = n \left( \frac{h}{2\pi} \right)$$

$\frac{L_1}{L_2} = \frac{n_1}{n_2}$

check  $m \left( 2.18 \times 10^6 \frac{Z}{n} \right) \left( 0.529 \times \frac{n^2}{Z} \times 10^{-10} \right)$

# independent of ( $Z$ )

# linear momentum ( $P = mv$ )  $\Rightarrow P = m \left( 2.18 \times 10^6 \frac{Z}{n} \right)$

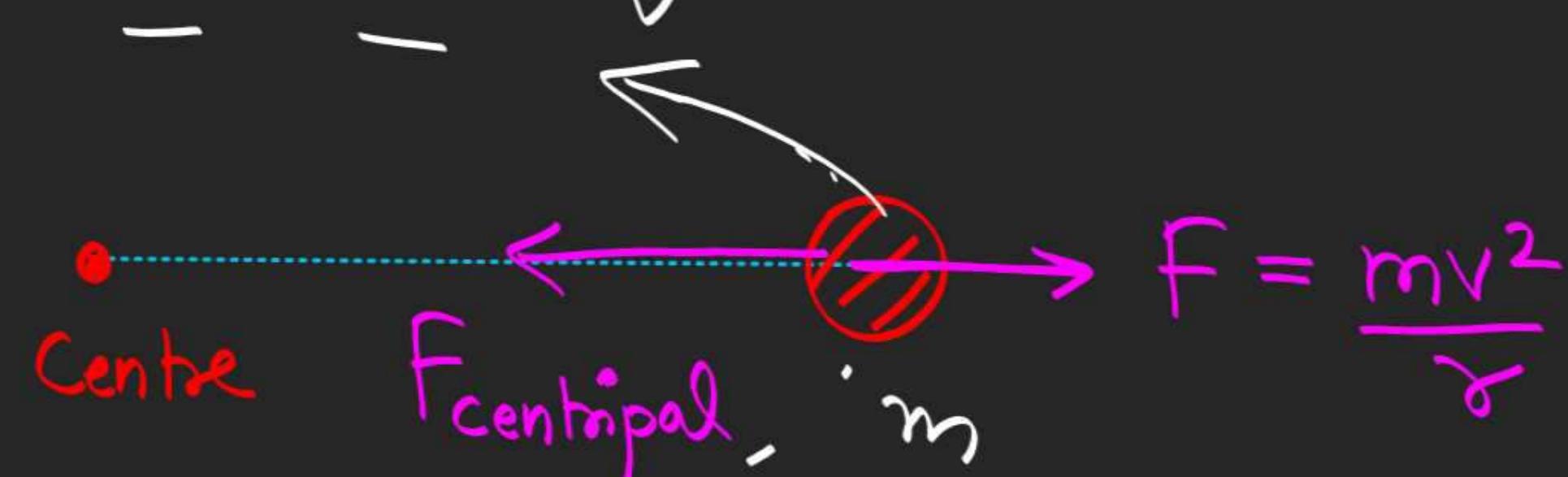
$(P \propto \frac{Z}{n})$

Centrifugal force

$$F = \frac{mv^2}{r}$$

$$\Rightarrow F \propto \frac{v^2}{r}$$

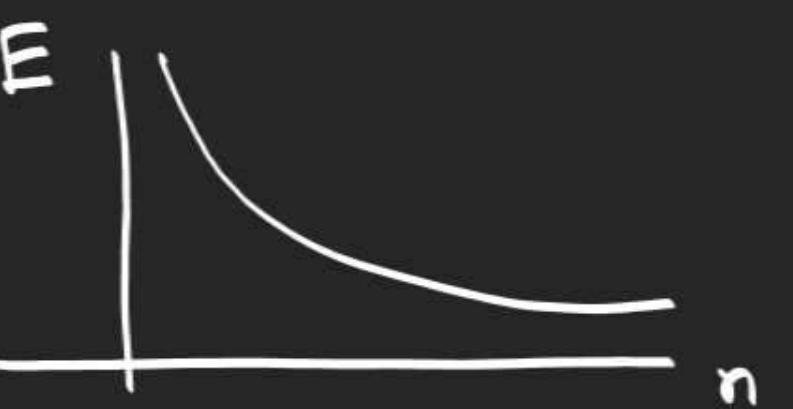
$$F \propto \frac{(z/n)^2}{\left(\frac{n^2}{z}\right)}$$



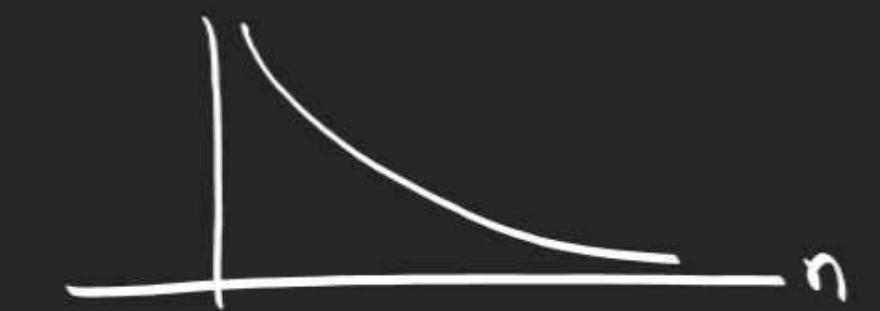
$$F = \frac{mv^2}{r}$$

$$\boxed{F \propto \frac{z^3}{n^4}}$$

(i)  $E_n \propto \frac{z^2}{n^2}$



(ii)  $v \propto \frac{z}{n}$



(iii)  $\gamma \propto \frac{n^2}{z}$



(iv)  $T \propto \frac{n^3}{z^2}$



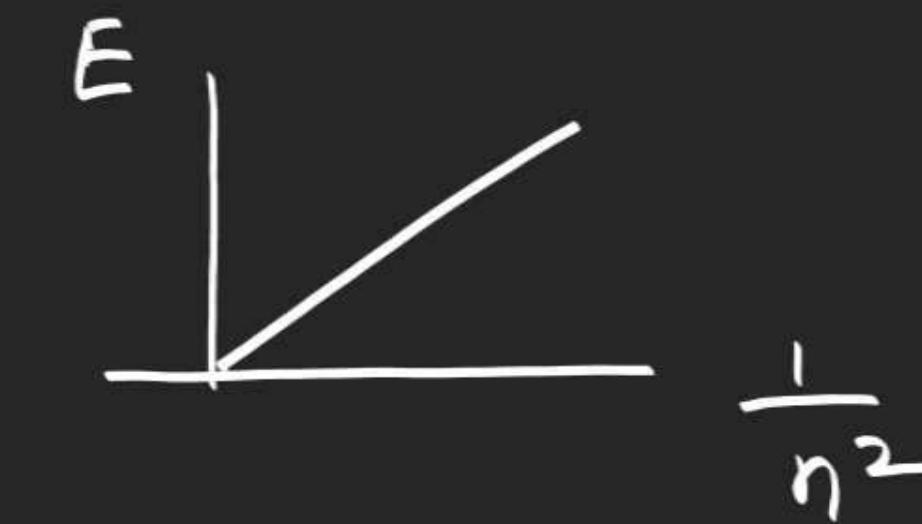
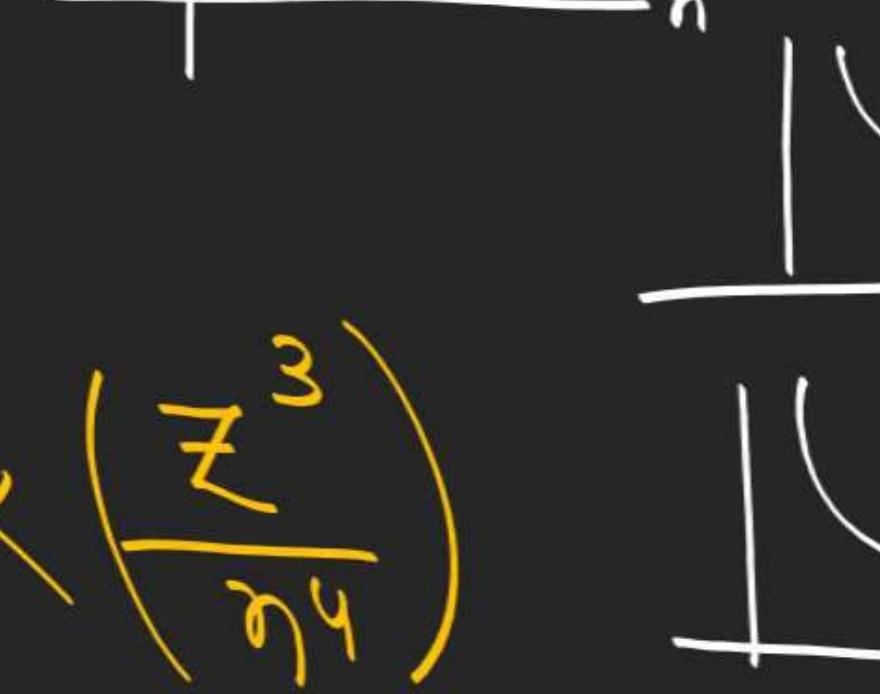
(v)  $\overline{L} \propto (n)$



(vi)  $\overline{P} \propto z/n$



(vii)  $\overline{F}_{\text{centrifugal}} \propto \left(\frac{z^3}{n^4}\right)$



Spectral emission : When e- from higher level to lower level transition

then rad<sup>n</sup> are emitted.

⇒ studied by Rydberg

other atom

$$\frac{1}{\lambda} = RZ^2 \left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$$

⇒ wave no. ⇒  
(for H-atom)

$$\frac{1}{\lambda} = R \left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right] \text{ (cm)}$$

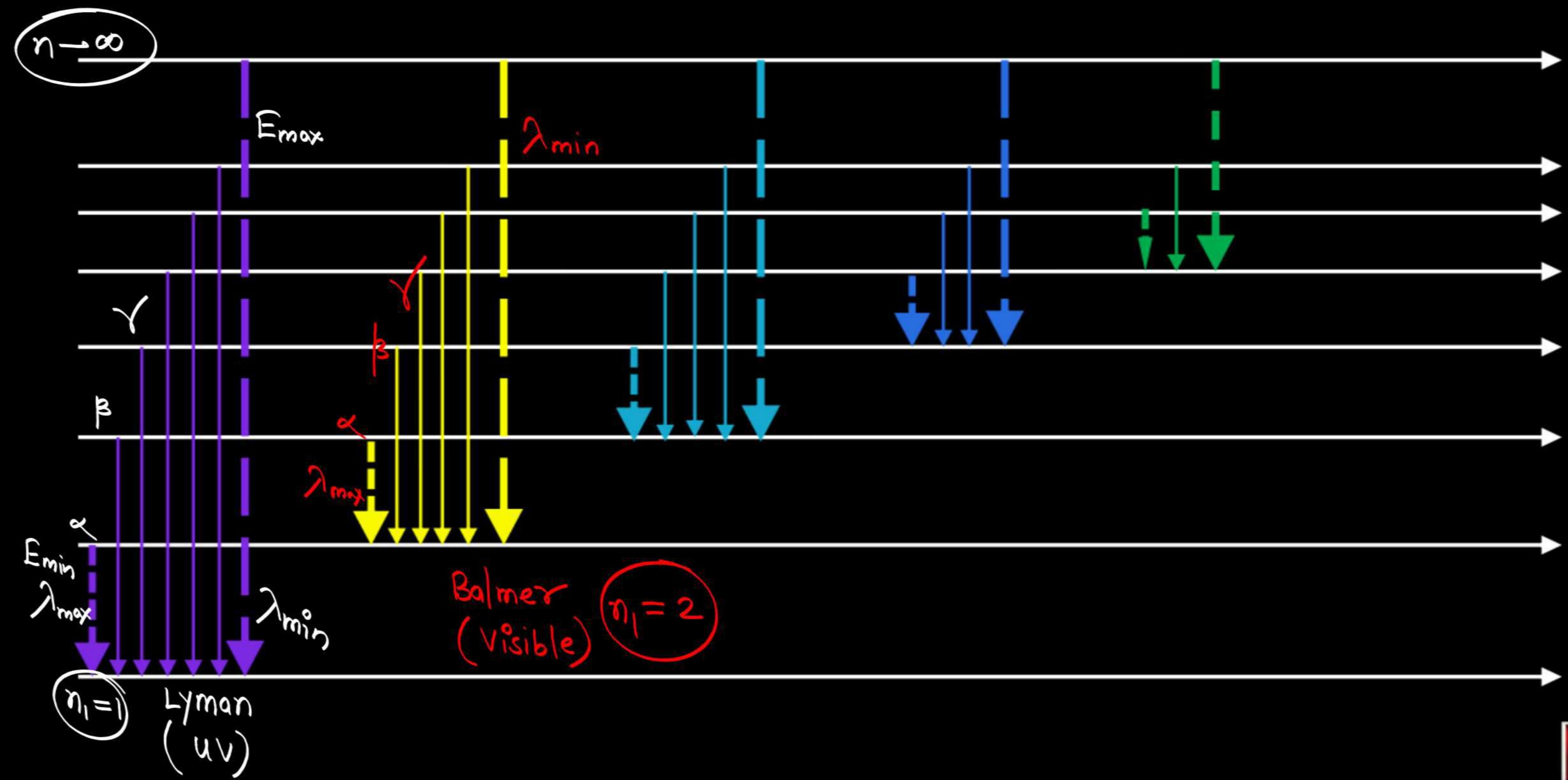
where  $n_2 \rightarrow$  Higher level  
 $n_1 \rightarrow$  lower level

Various series  
were discovered

- $n_1 = 1 \Rightarrow$  Lyman Series (uv rays)
- $n_1 = 2 \Rightarrow$  Balmer series (visible)
- $n_1 = 3 \Rightarrow$  Paschen series
- $n_1 = 4 \Rightarrow$  Brackett series
- $n_1 = 5 \Rightarrow$  Pfund series

} infrared

# The Spectral Lines for Atomic Hydrogen



LIVE

## Lyman series (Analysis)

I From  $n_2 = 2, 3, 4, \dots, \infty$  to  $n_1 = 1$

II fall under (uv) range

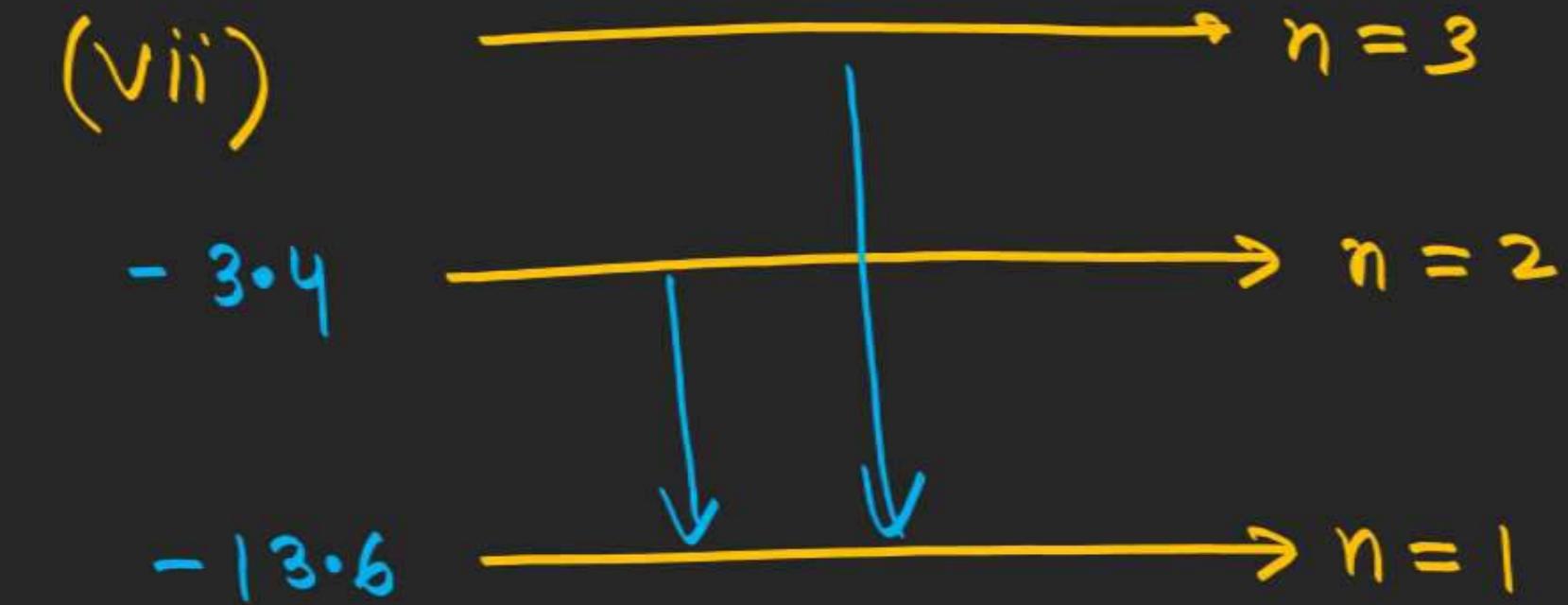
III  $\frac{1}{\lambda} = R \left[ \frac{1}{1^2} - \frac{1}{n_2^2} \right]$

(iv)  $\frac{1}{\lambda_{\max}} = R \left[ \frac{1}{1^2} - \frac{1}{2^2} \right] \Rightarrow \lambda_{\max} = \frac{4}{3R}$

(v)  $\frac{1}{\lambda_{\min}} = R \left[ \frac{1}{1^2} - \frac{1}{\infty^2} \right] \Rightarrow \lambda_{\min} = \frac{1}{R}$

(vi)  $\frac{\lambda_{\max}}{\lambda_{\min}} = \frac{4}{3}$

$\lambda_{\max}$  neighbour  
 $\lambda_{\min}$  ( $\infty$ )



Photon |  $\text{rad}^{-1}$  |  $e^-$  energy

$$E \geq 10.2 \text{ eV}$$

$$\lambda \leq 121.56 \text{ nm}$$

## Numerical

$$\frac{\lambda_1}{\lambda_2} = \frac{\left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]_2}{\left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]_1}$$

(find)  $\rightarrow \frac{(\lambda_{\max})_{\text{Paschen}}}{(\lambda_{\min})_{\text{Lyman}}} = \frac{\left( \frac{1}{1^2} - \frac{1}{\infty^2} \right)}{\left( \frac{1}{3^2} - \frac{1}{4^2} \right)} = \left( \frac{144}{7} \right)$

$$\frac{\lambda_{\max}}{\lambda_{\min}} = \frac{\left[ \frac{1}{n_1^2} - \frac{1}{\infty^2} \right]_{\min}}{\left[ \frac{1}{n_1^2} - \frac{1}{(\text{पड़ोसी})^2} \right]_{\max}}$$

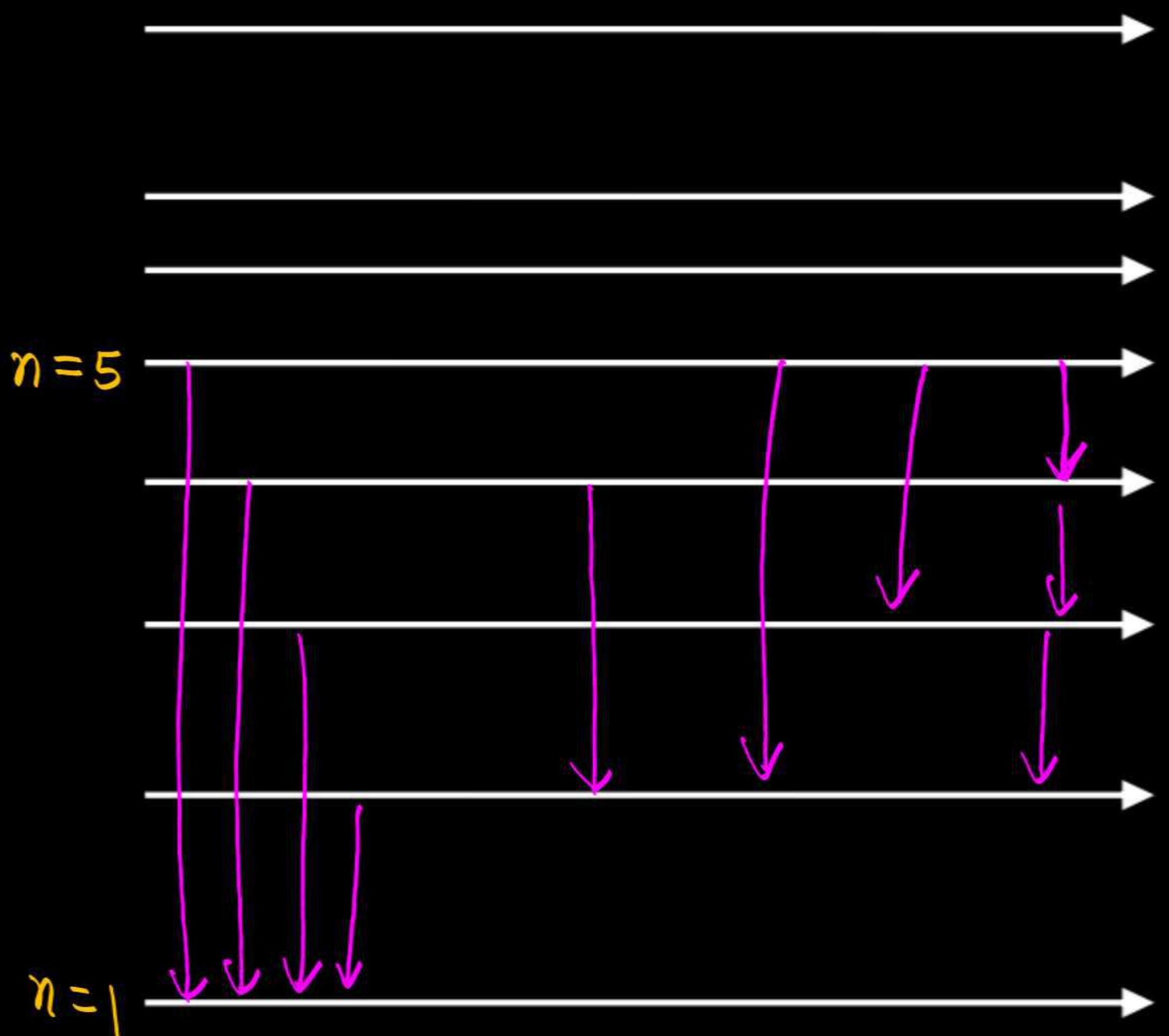
Balmer  $\rightarrow \frac{\lambda_{\max}}{\lambda_{\min}} = \frac{\left( \frac{1}{2^2} - \frac{1}{\infty^2} \right)_{\min}}{\left( \frac{1}{2^2} - \frac{1}{3^2} \right)_{\max}}$

$= \left( \frac{9}{5} \right)$

No. of spectral lines

(i) Total lines =  $\frac{(\Delta n)(\Delta n + 1)}{2}$

example from  $n = 5 \text{ to } 1$



(i) Total lines =  $\frac{4 \times 5}{2} = 10 \text{ lines}$

$\frac{\text{Lyman (UV)}}{(n_2 - 1)}$        $\frac{\text{Balmer (Visible)}}{(n_2 - 2)}$        $\frac{\text{Infrared}}{3}$

= ④                                  (3)

LIVE

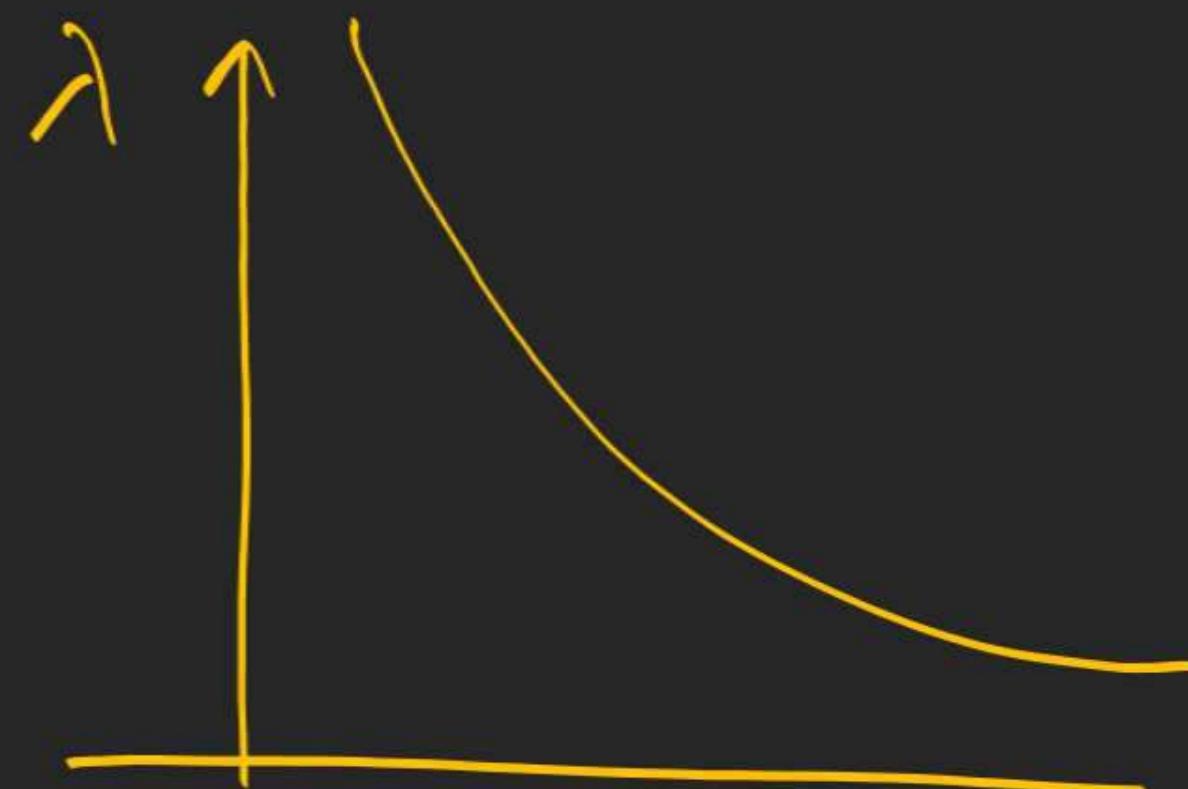
## Dual Behaviour of matter : De-Broglie (Hypothesis)

Every moving object possesses dual behaviour

and wavelength associated is called De-Broglie wavelength.

Mathematically

$$\lambda = \frac{h}{P} = \frac{h}{mv}$$



(P) or (v)

where  $m$  = mass of particle

$v$  = velocity —

$$(*) \lambda = \frac{h}{mv} = \frac{h}{p} = \frac{h}{\sqrt{2mK}} = \frac{h}{\sqrt{2mqV}} \quad \text{velocity} \quad \text{Potential diff}$$

# For  $e^-$  :  $\frac{h}{\sqrt{2meV}} = \sqrt{\frac{150}{V}} A^\circ = \frac{12.2}{\sqrt{V}}$

Comparison  $\Rightarrow$

$$(i) \frac{\lambda_1}{\lambda_2} = \frac{P_2}{P_1} = \frac{m_2 v_2}{m_1 v_1} = \sqrt{\frac{m_2 K_2}{m_1 K_1}} = \sqrt{\frac{m_2 q_2 V_2}{m_1 q_1 V_1}}$$

## DE BROGLIE WAVELENGTH

according to de Broglie,  
every object in motion  
has a  
wave character

- ✓ For different charge particle put value of (q)
- ✓ For electron e=  $1.6 \times 10^{-19}$
- ✓ For  $\alpha$  particle put 2e
- ✓ For proton put (e)

$$\lambda = \frac{h}{mv}$$

$$= \frac{h}{p}$$

$$= \frac{h}{\sqrt{2m(K.E.)}}$$

$$= \frac{h}{\sqrt{2m(qV)}}$$

$$= \sqrt{\frac{150}{V}} A^\circ$$

$$= \frac{12.2}{\sqrt{V}} A^\circ$$

LIVE

## De-Broglie's comparison with Bohr model

#  $\lambda = \frac{h}{mv}$  and  $mv\gamma = n \frac{h}{2\pi}$

$$2\pi\gamma = n \left( \frac{h}{mv} \right)$$

#  $n^{th}$  orbit :  $(n)\lambda$  produced

$$2\pi\gamma = n\lambda$$

...> (circumference is  
Quantised

#  $2\pi \left( 0.529 \frac{n^2}{Z} \right) = n\lambda$  ←-----

⇒ means integral multiple  
of  $(\lambda)$

⇒  $\left( \frac{10}{3} \right) \frac{n}{Z} = \lambda$  ⇒ De-Broglie wavelength  
in  $n^{th}$  orbit

$$1. E_n = -13.6 \frac{Z^2}{n^2}$$

$$2. KE = |E_n| = -\frac{P \cdot E_0}{2}$$

$$3. (n) \uparrow \quad KE \downarrow \quad (E, PE) \uparrow$$

$$4. \Delta E = -13.6 \left( \frac{1}{n_1^2} - \frac{1}{n_2^2} \right)$$

$$5. \text{ Same energy / Same } (\lambda) \text{ or photon}$$

H	$\text{He}^+$	$\text{Li}^{+2}$
$n$	$2n$	$3n$

$$6. \frac{\text{Energy Gap}}{\text{Same } (n)}$$

H	$\text{He}^+$	$\text{Li}^{+2}$
$\propto$	$4x$	$9x$

$$7. \Delta E = \frac{hc}{\lambda} = \frac{1240}{\lambda \text{ (nm)}}$$

$$8. \frac{1}{\lambda} = R Z^2 \left[ \frac{1}{n_1^2} - \frac{1}{n_2^2} \right]$$

$$9. \frac{\lambda_1}{\lambda_2} = \left( \frac{\text{?}}{\text{?}} \right)^2$$

$$10. mv\tau = n \frac{h}{2\pi}$$

$$11. \gamma_i = 0.529 \frac{n^2}{Z}$$

$$12. V_n = 2.018 \times 10^6 \left( \frac{Z}{n} \right)$$

$$13) \lambda = \frac{h}{mv}$$

$$\Rightarrow 2\pi r = n\lambda$$

$$14) \lambda = \left( \frac{10}{3} \right) \frac{n}{Z}$$

$$15) F \propto \frac{Z^3}{n^4}$$

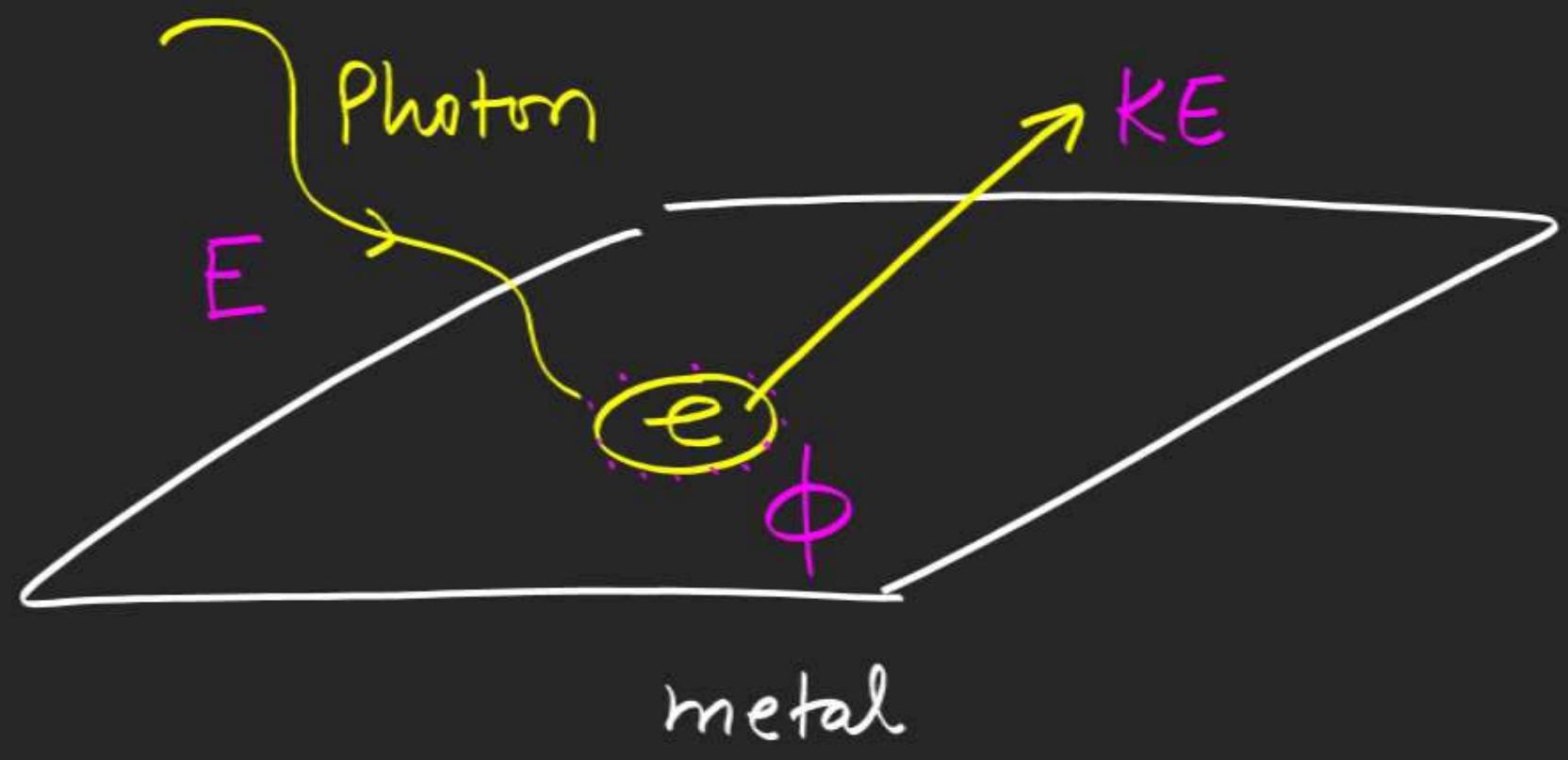
$$16) T \propto n^3/Z^2$$

$$17) f \propto Z^2/n^3$$

$$18) P \propto Z/n$$

$$19) F \propto -\frac{dV}{dr}$$

Photoelectric effect : Ejection of  $e^-$  from metal surface when photon / grad<sup>n</sup> of sufficient energy / freq / wavelength falls on it



# Min energy required is work function ( $\phi$ )

# Energy  $\geq \phi$  ..... ( $\phi$ ) is diff. for diff. metal

#  $\phi = h\nu_0$  when  $\nu_0$  = threshold / min freq

above which

Photoelectric effect

$$\# \phi = \frac{hc}{\lambda_0}$$

Cond<sup>n</sup> for photoelectric effect

- a)  $E \geq \phi$
- b)  $\nu > \nu_0$
- c)  $\lambda \leq \lambda_0$

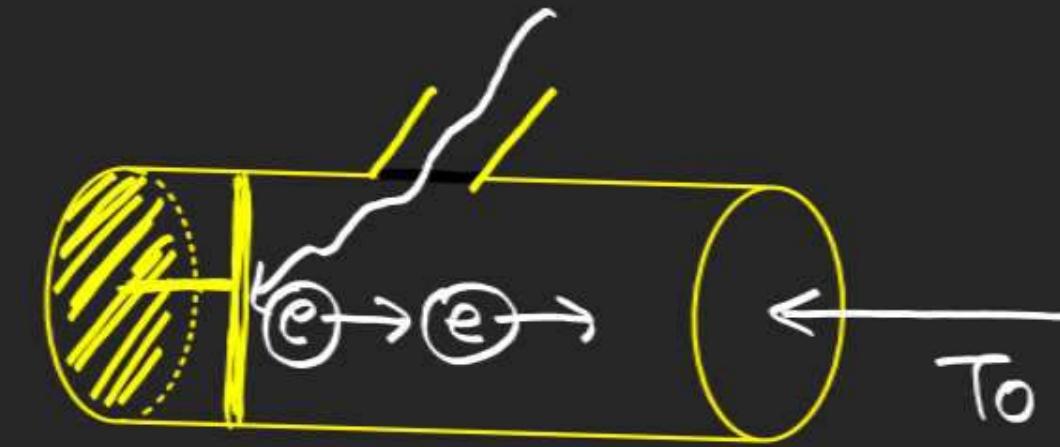
occurs  
when  $\lambda_0$  = threshold / max wavelength  
Below which photoelectric effect occurs

## Important points

- 1.) one photon interact with one  $e^-$
  - 2.) Instantaneous process (no time lag)
  - 3.) Elastic collision (no energy loss)
  - 4.)  $E = KE + \phi$
  - 5.) ( $K.E.$ ) of  $e^-$  depends only on freq / wavelength of rad/ $\gamma$  / photon
  - 6.)  $KE$  is independent of intensity
  - 7.) Current produced is called photocurrent
  - 8.) Current  $\propto$  Intensity
  - 9.)
- # Intensity = no. of photon/sec.
- To stop these  $e^-$  opp. potential diff is req. is called stopping Potential

$$\Rightarrow KE = qV_s$$

$$\Rightarrow KE = eV_s$$



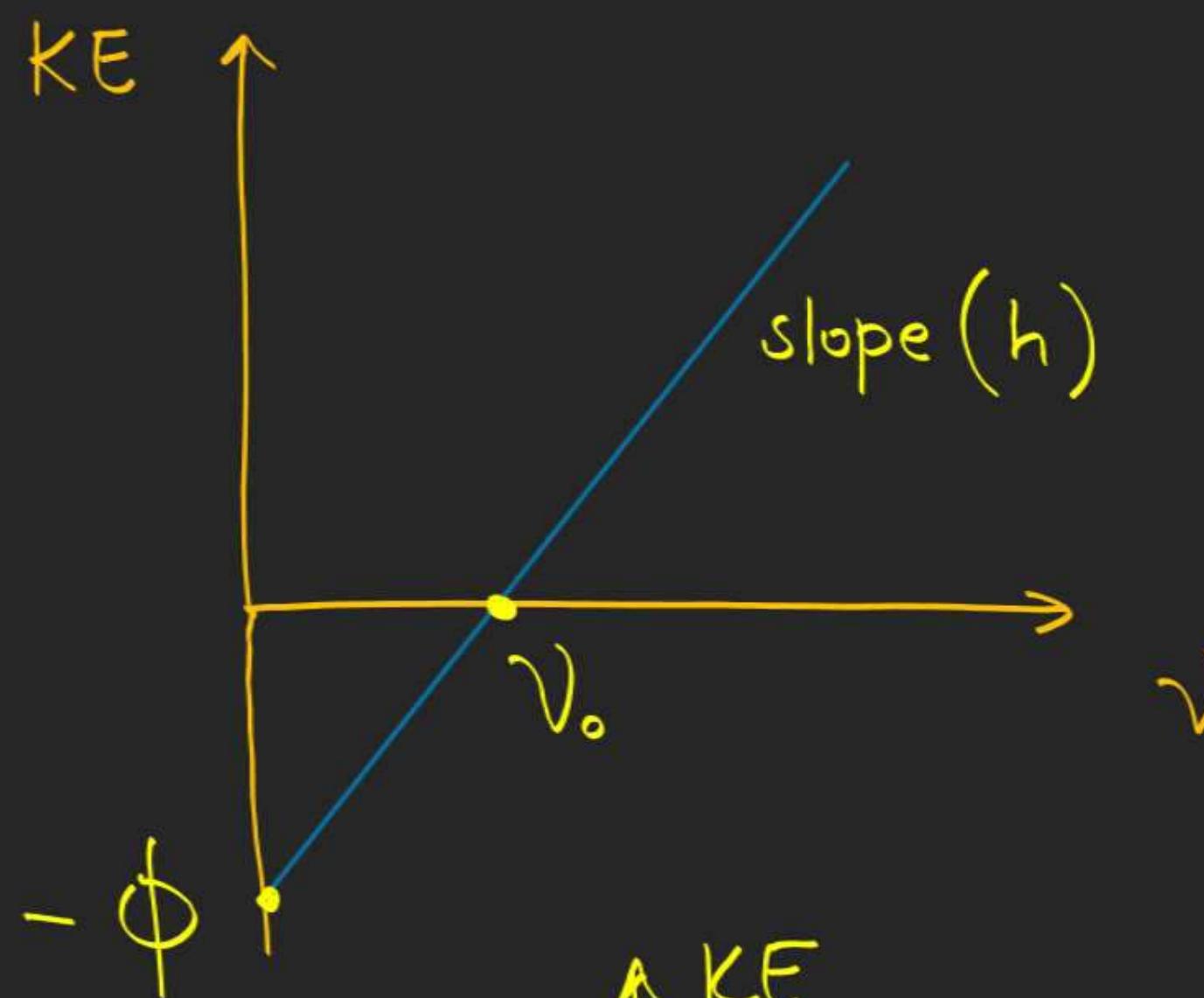
## Numericals

$$1. \quad E = KE + \phi$$

$$2. \quad KE = E - \phi$$

$$\Rightarrow KE = h\nu - h\nu_0$$

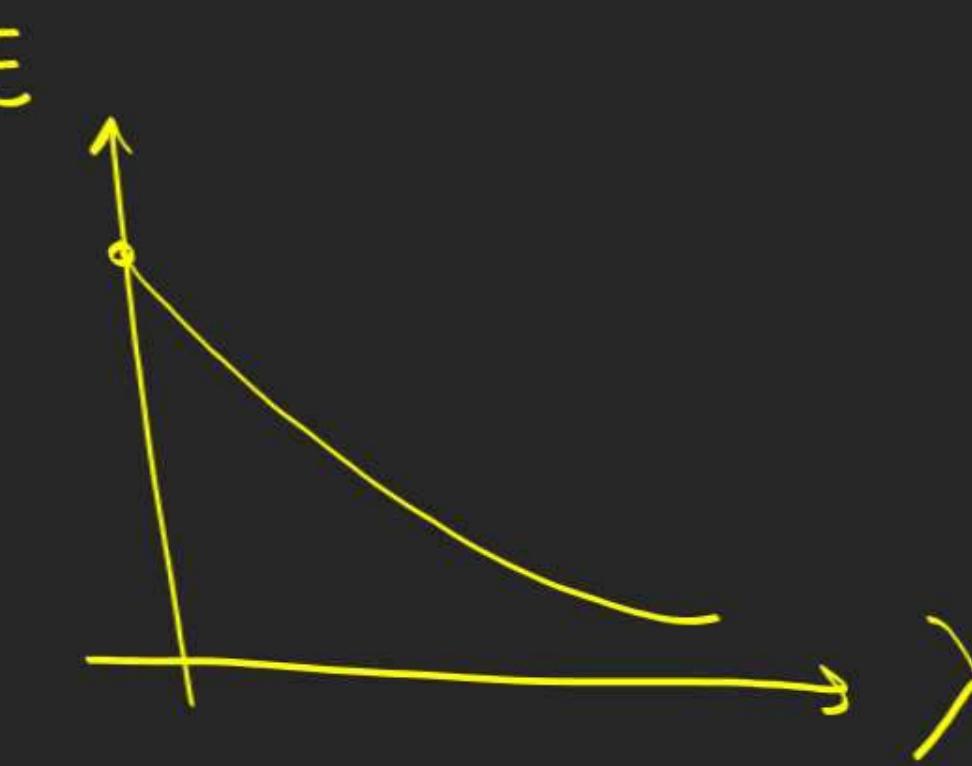
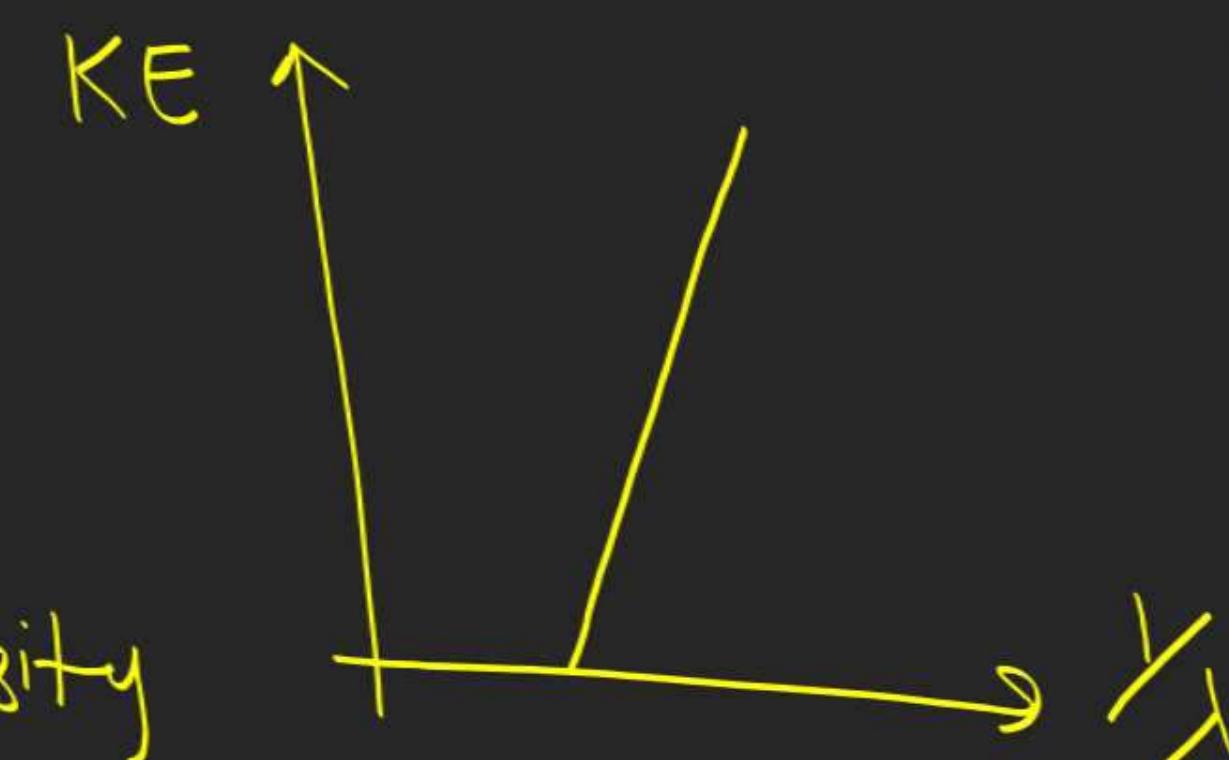
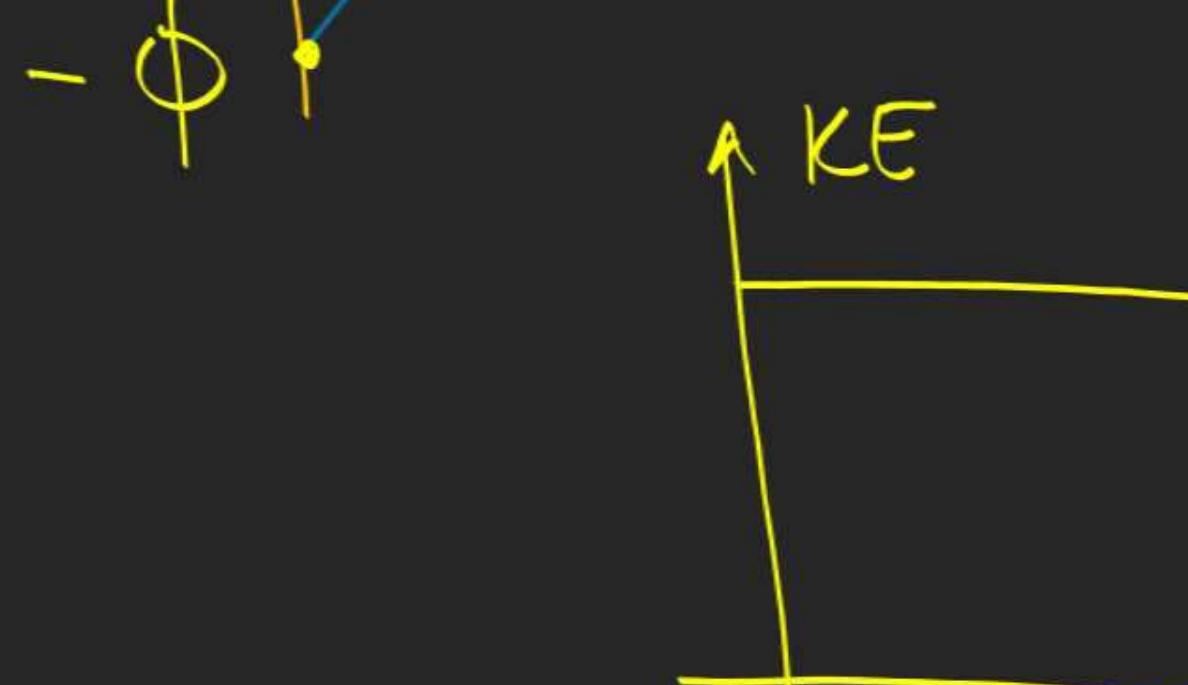
$$\Rightarrow KE = \frac{hc}{\lambda} - \frac{hc}{\lambda_0}$$

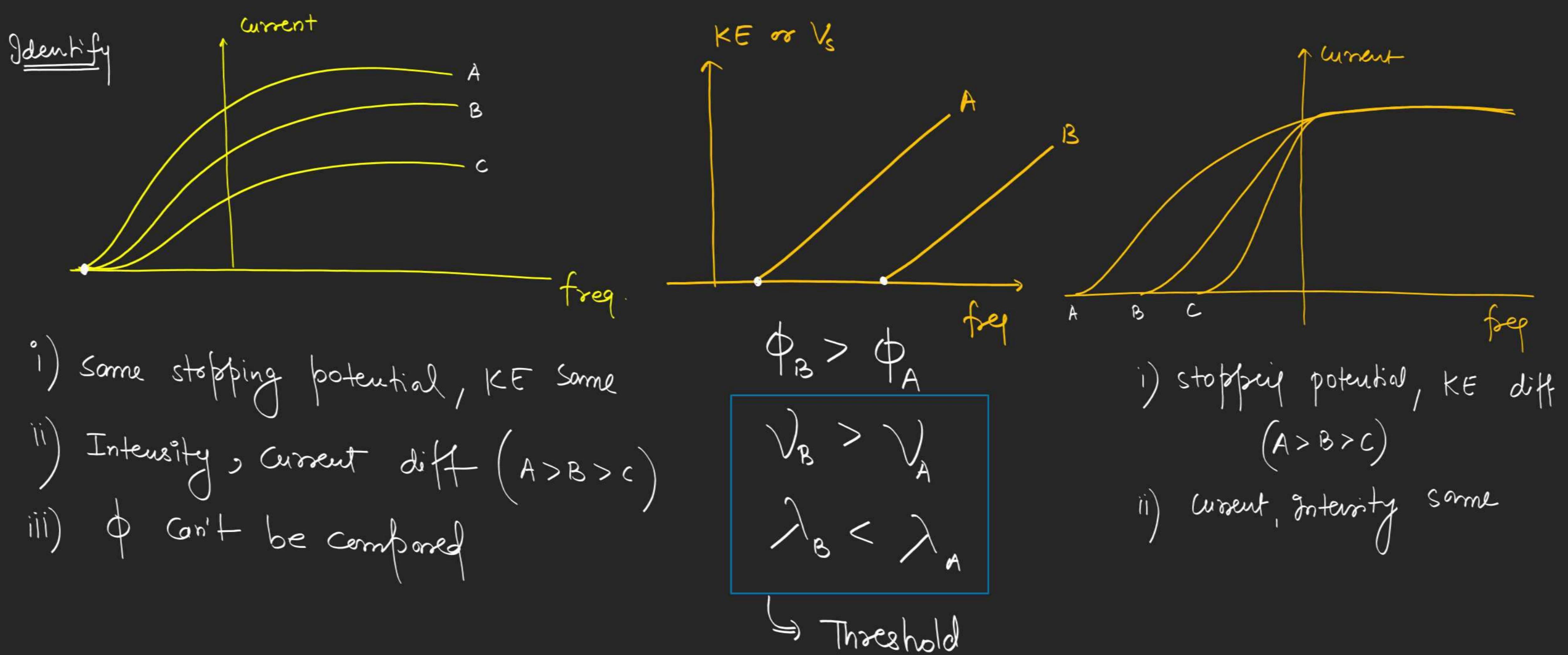


$$\Rightarrow eV_s = h\nu - h\nu_0$$

$$\Rightarrow V_s = \frac{h}{e}\nu - \frac{h}{e}\nu_0$$

$$\Rightarrow V_s = \frac{hc}{e\lambda} - \frac{hc}{e\lambda_0}$$





## Heisenberg's Uncertainty Principle

It states that it is impossible to determine simultaneously, the exact position and exact momentum (or velocity) of an electron.

# There is uncertainty in value of position ( $\Delta x$ ) & momentum ( $\Delta p$ ) of e-

$$\# \quad \Delta x \cdot \Delta p \geq \frac{h}{4\pi}$$

$\Rightarrow$  If error in velocity ( $\Delta v$ )

$$\Rightarrow \Delta x (m \Delta v) \geq \frac{h}{4\pi}$$

$$\Delta v \geq \frac{h}{4\pi m \cdot \Delta x}$$

LIVE